DARK SIDE

# DarkSide

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### The DarkSide project

Dual phase Liquid Argon TPC to search for Dark Matter

#### DarkSide-50 (2013-2018)

- Nested detectors: 50 kg Liquid Argon TPC Boron-loaded Liquid Scintillator Veto Water Cherenkov Veto \_\_\_\_\_\_
- First use of underground argon (depleted of <sup>39</sup>Ar)
  - Demonstrate pulse shape discrimination (>10<sup>7</sup>)
- DarkSide-20k (2020-2026)
  - 30 ton (20 ton FV) depleted Liquid Argon TPC
  - Designed to be background-free at 100 t yr exposure
  - Similar system of vetoes as DS-50, use of SiPM
- **ARGO** (2025-2035)
  - 300 ton (200 ton FV)
  - Search for WIMPs up to the neutrino floor
  - Solar neutrino physics



#### **DarkSide Collaboration**

Large international collaboration (43 institutes, 277 signatures for the DS-20k proposal)

Entire Liquid Argon community for DM searches to join this project → no competition with Liquid Argon for DS-20k

Part of the collaboration works on external calibration experiments, using neutron beams to study the response of LAr to Nuclear Recoils

- SCENE in the US (2014)
- ARIS experiment at IPNO (2016)

Future experiments including directionality (ReD, ARIS-2)

#### IN2P3 DarkSide groups

**DS 1ton prototype** 

#### **DS-50** analyses

3 papers in preparation (G4DS Monte Carlo, f90 (PSD) model, DS-50 background budget)

#### **DS-20k optimization**

(do we need vetoes to be background-free? yes! DS-20k sensitivity and discovery potential, solar neutrinos, ...) APC+LPNHE Developers of the Monte Carlo simulation NR calibration experiments and directionality ARIS (done) ARIS-2 ReD

ARIA

(depletion factor needed for DS-20K and ARGO)

French groups in DS have leading roles in physics analyses
 Leaders of the DS MC simulation
 External calibration with neutrons beam



	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33
DS-50																		
DS-20k																		
ARGO																		

- DarkSide-50 (37 kg FM) will take data at LNGS until 2018
  DarkSide-20k (20 ton FM) will start data taking at LNGS in 2021
  - Expect approval from INFN and NSF in April 2017
  - A 1-ton scale prototype will be built starting next year to test the technologies needed for DS-20k
  - ARGO will follow DS-20k (200 ton FM)

# What is DarkSide-50?

a two phase liquid argon (LAr) detector, within a neutron veto, within a muon veto, under a mountain

#### . and all because of backgrounds

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### DarkSide-50

- Experiment installed at LNGS
- Double phase TPC with 50 kg of LAr
- Vetoes system: Liquid Scintillator and Water Cherenkov
- 2 data releases published
  - Atmospheric Ar campaign (52 days) Phys.Lett.B.743(456)
  - Underground Ar campaign (71 days) Phys. Rev. D 93, 081101 (2016)

Background reduction Depleted Underground Argon Low background materials Active Shields Background identification Pulse Shape Discrimination S1/S2 discrimination Measure neutron flux in borate scintillator

Demonstrate the potential of the technology for multi ton background-free detector

10 m (high) x 11 m (diameter) Water Tank



#### DarkSide-50 detector design

#### Liquid Argon TPC

36 cm x 18 cm radius 50 kg (36.9 fiducial) 38 PMTs Uniform electric field (200 V/cm) 1 cm gas pocket, extraction field (2.8 kV/cm) Reflectors and TPB coating Liquid Scintillator Veto (LSV) 30 tons, 2m radius LS (1:1 TMB+PC) 110 PMTs

Water Cherenkov (WCD)

1 kton water, 5.5m radius 80 PMTs



#### Ionization and excitation

- ➢ WIMPs scatter on LAr → primary photons (S1) + electrons drifting to the GAr region where they are accelerated and emit light through electroluminescence (S2)
  - S1/S2 ratio allow to distinguish electron recoils from nuclear recoils with a rejection power of  $\sim 10^2 10^3$





# S2/S1 Ratio

[arb]

380

2.2∟ 0



- Exploit different fractions of energy going to excitation (S1) and ionization (S2)
- Allow to distinguish ER from NR Singlet RTriplet with a discrimination power of 10<sup>2</sup>-10<sup>3</sup>
- Technique available for dual phases noble liquids detector (Xe or Ar)

Benetti et al. (ICARUS) 1993; Benetti et al. (WARP) 2006



161301 (2016)

S1 (phd)

30

40

50

20

10

10

Er

### **Pulse Shape Discrimination in LAr**



### **Atmospheric Argon results**

50 days of data taking

1.5x10<sup>7</sup> single scatter events in the ROI (dominated by <sup>39</sup>Ar)
 None of them enter the WIMP search region



## **Underground Argon**

- Atmospheric Argon contains cosmogenic contamination from <sup>39</sup>Ar due to activation from cosmic rays
  β-decay with a rate of ~1 Hz/kg in Atmospheric Argon
  Half-life of <sup>39</sup>Ar is 270 years
  Solution: use underground argon naturally depleted of <sup>39</sup>Ar
  - DS-50: 150 kg of UAr extracted from a mine in Colorado, purified at FNAL and shipped to LNGS







#### **Underground Argon run**



#### **DS MC simulation: G4DS**

- The MC simulation of DS has been written by French groups
- GEANT4 MC simulation including all the geometries of the DS program (DS-50, DS-20k, DS-1ton, ARIS, ReD...)
  - Optical tuning based on DS-50 data
- Fundamental tool for the analysis of DS-50 data and the design of DS-20k

DarkSide-50 Geometry







DGEU 4.





# darkish3D fit to extract <sup>39</sup>Ar component

e<sup>22000</sup>



#### <sup>39</sup>Ar depletion factor

<sup>39</sup>Ar = 0.71±0.05 mBq/kg → depletion factor of 1400±200 (previous upper limit was >100) Discovery of the <sup>85</sup>Kr component ~ 2 mBq/kg (later confirmed by coincidence search)



### Background-free 52 days exp (AAr)



#### Background-free 70 days exp (UAr)



#### 70 day exp (UAr) + S1/S2 cut



### DarkSide-20k

- Dual phase TPC → from 50 kg to 30 ton (20 ton FV)
- TPC light readout with SiPM
  - Better light yield (>10 pe/keV)
  - Cleaner than PMTs and lower mass (much lower neutron background)
- Active vetoes (LS and WT) similar as DS-50
- Depleted Argon: extract large quantities of UAr in Colorado (URANIA) and further purify it with a distillation column (ARIA) already partially funded
  - Scientific goal: 100 ton yr background-free exposure



#### Yellow Book

The DarkSide-20k Yellow Book Technical Proposal/Pre-Technical Design Report September 5<sup>th</sup>, 2016



#### 43 institutes 277 signatures

Submitted to the LNGS Scientific Committee and to INFN/NSF

Joint review started → expect approval by April 2017

#### SiPM R&D

- The other big challenge of DS-20k is to instrument a large area of detector with SiPM (total active surface of ~15 m<sup>2</sup>)
- A lot of R&D is on-going mainly driven by LNGS and FBK
- Already reach goal for Dark Count rate and photon-detection efficiency







### **Underground Argon procurement**

A: plant expansion to extract 100 kg/day of UA Colorado

ARIA: ~300 m distillation column to be installed at Seruci (Sardinia) to remove  ${}^{85}$ Kr from Ar (easy) and to further reduce the residual  ${}^{39}$ Ar component ( $\rightarrow$  factor of 10 per pass)

These programs have already been funded





First ARIA module OCERN

#### DarkSide-20k physics goal



#### **ARIS** experiment



- Experiment performed at IPNO using the LICORNE beam
- Use a small dual phase TPC built at UCLA
- Installation of the detector on the LICORNE beam line in September 2016
- Data taking with the neutron beam October 3-14
- Analysis of the data is on-going

#### What is **ARIS**

Expose te TPC on the highly collimated neutron beam RGON RECOIL IONIZATION AND SCINTILLATION (ARIS) Neutrons with mean energy of 1.44 MeV, beam opening O. Riffard', B. Schiltzerf, Y. Strong, A. Terezol' H. Watel V. Watel J. Wilson<sup>2</sup> C. Riffard', B. Schiltzerf, Y. Strong, A. Terezol' H. Watel V. Watel J. Wilson<sup>2</sup> Laboration of ~6 degrees → thanks to the neutron production

through inverse kinematics



Neutrons scatter in the TPC and are observed with EDEN detectors at different angles (i.e. different recoil energies)

Goal: characterize LAr response for scintillation, ionization and PSD as function of the recoil energy



## **Experimental setup**



### **ARIS** experimental setup

- We used 8 EDEN neutron detectors (energy from 7 keV to 120 keV)
- Beam energy as small as possible to maximize collimation
- 5 different electric fields in the TPC (from 0 to 500 V/cm)







- Pulsed neutron beam
- TPC instrumented with 8 PMTs
  - ◆ 7 I-inch PMTs on the top
  - + 1 3-inch PMT on the bottom
- + EDEN neutron detectors for:
  - + TOF
  - + n/γ PSD
- Triple coincidence between beam, TPC and ND was requested

#### **ARIS first results**

#### <sup>241</sup>Am 60 keV γ peak LY~7 pe/keV



#### **TPC/Beam TOF vs f90**



#### Importance of IPNO beam

- LICORNE is a directional neutron source at the Tandem accelerator at IPNO
- <sup>7</sup>Li is accelerated to 13-17 MeV and neutrons are produced through the reaction p(<sup>7</sup>Li,n)<sup>7</sup>Be
  - High fluxes of neutrons (up to 10<sup>8</sup> n/s/sr)
  - Highly collimated (from ~2 to ~25 degrees depending on the Li energy)
- This is a unique facility that can be exploited to characterize TPC response for low energy nuclear recoils
- The success of ARIS experiment demonstrated that this facility can be used for the DS external calibration program
- In-kind contribution that IN2P3 can offer to DarkSide

#### **External calibration program**

- The DarkSide collaboration is involved in an external calibration program for the characterization of the LAr response to low energy to NR
- SCENE experiment @ Notre Dame → isotropic neutron beam
- ARIS experiment @ IPNO → highly collimated and monoenergetic neutron beam
- Next: ReD: an experiment to sense directionality in LAR



- TPC built at Naples, instrumented with SiPM (2 5x5cm arrays on top and bottom)
- Minimize materials out from LAr active region to maximize neutrons single scatter in LAr
- Expose it to a neutron beam

LICORNE@IPNO is a perfect candidate

### Directionality





as in Case 1, there will be more electron-ion recombination since the electron passes more ions as it drifts through the chamber.

- Columnar recombination: more S1 and less
  S2 if track parallel to electric field
- Exploit the preferred direction of WIMPs with respect to isotropic v and n backgrounds
- Might allow to go beyond the neutrino floor
- Some hints from SCENE but need more precise measurements on neutron beams with dedicated setup (ReD at IPNO)



35

More Recombination

#### ARGO

DarkSide-20k will be followed by ARGO (Argon Observatory)
 300 ton Dual Phase TPC to be operational in 2026
 Search for WIMPs through the neutrino floor
 Do solar neutrinos physics



#### Solar neutrinos with ARGO

#### D. Franco et al, JCAP08(2016)017

- Excellent scintillator 40 photons/keV
- Easy to purify, high intrinsic radiopurity wrt organic liquid scintillators
- Hundreds of ton of mass
- Exceptional PSD

400 ton yr exposure: First observation of <u>CNO neutrinos</u> measure <sup>7</sup>Be-ν (~2%), <u>pep-ν (~10%)</u>



#### Conclusions

- DarkSide has a 20 years physics program built on the success of DarkSide-50
  - Excellent PSD of LAr, establish the use of UAr for DM searches
  - Already proved a background-free exposure equivalent to 5.5 ton yr
- We have submitted a pre-TDR to INFN and NSF to propose Darkside-20k@LNGS
  - Largest DM collaboration (unify the entire Argon community)
  - Background-free exposure of 100 ton yr → limit on  $\sigma_{SI}$ ~1.2x10<sup>-47</sup> cm<sup>2</sup> (1.1x10<sup>-46</sup> cm<sup>2</sup>) for 1 TeV(10 TeV) WIMPs
  - Expect approval from INFN and NSF in April 2017
  - French groups have a leading role in the physics of DarkSide
    - Analyses of DS-50

Design optimization of DS-20k

The ARIS experiment demonstrate the availability in France of LICORNE@IPNO as a unique facility for the DS external calibration program that will be performed in the next years