# Mass Composition Studies with the Pierre Auger Observatory

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photo by S.J. Saffi, University of Adelaide

# **The Pierre Auger Observatory**



#### Air Shower Detection with the Pierre Auger Observatory



# **Energy Spectrum of UHECRs**



exposure at UHE:  $(5.34 \pm 0.13) \times 10^4$  km<sup>2</sup> sr yr

#### **Mass Composition Studies**

# (a) Longitudinal Development of Air Showers



# Average Shower Maximum vs. Energy



 $\langle X_{
m max}
angle \propto D_{
m 10} \log{(E/A)}$  (mass A, energy E, elongation rate  $_{
m 10}$   $\sim$  54 - 64 g/cm²/decade)

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# Standard Deviation of $X_{\text{max}}$

• 
$$\sigma(X_{\max})^2_A = \lambda^2_A + \sigma(X_{\max} - X_{\text{first}})^2_A$$

$$\ \, \bullet \ \, \sigma(X_{\max})_{\rho} > \sigma(X_{\max})_{A} > \sigma(X_{\max})_{\rho}/\sqrt{A}$$

mixed composition:

$$\sigma(X_{\max})^2 = \langle \sigma_i^2 \rangle + \left( \left\langle \langle X_{\max} \rangle_i^2 \right\rangle - \langle X_{\max} \rangle^2 \right)$$



# Standard Deviation of X<sub>max</sub> vs. Energy



# Standard Deviation of $X_{max}$ vs. Energy



# Average Shower Maximum: Comparison to Telescope Array



Auger-TA Working Group on Composition (UHECR conference series)



UHECR16, Kyoto



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#### **Comparison to Telescope Array**



 $\langle \Delta \rangle = (2.9 \pm 2.7 \text{ (stat.)} \pm 18 \text{ (syst.)}) \text{ g/cm}^2$ 

# Fit of X<sub>max</sub> Distributions



Pierre Auger Coll., PRD 90 (2014) 12, 122006

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# **Mass Composition Studies**

# (b) Correlation of $X_{\text{max}}$ and Ground Signal



#### **Correlation of Xmax and Ground Signal**

 $18.5 < \lg(E/eV) < 19.0, X_{max}^*/S^*(1000)$ : scaled to  $10^{19} eV$ 



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#### **Correlation of Xmax and Ground Signal**

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# Mass Composition Studies (c) Surface Detector Data



# **Muons in Air Showers**



R. Ulrich, APS 2010

- muons from  $\pi^{\pm}$  decay at late stage of cascade ( $\lambda_{dec} \sim \lambda_{int}$ )
  - $\rightarrow\,$  number of generations  $\sim$  6 at 10^{19} eV
  - $\rightarrow\,$  amplified sensitivity to hadronic interactions
- X<sub>max</sub> is dominated by first interaction

### **Muon Production Depth**



Pierre Auger Coll., PRD D90 (2014) 1, 012012, erratum PRD92 (2015) no.1, 019903

# **Muon Deficit in Models?**







#### Summary: Mass Estimates with X<sub>max</sub>



# **Outlook: Upgrade of the Pierre Auger Observatory**

#### additional scintillators (4 m<sup>2</sup>)



 $X_{\text{max}}$  determination:



#### muon determination:



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