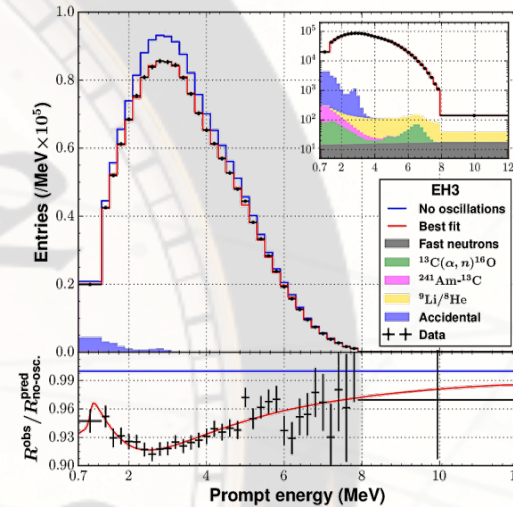
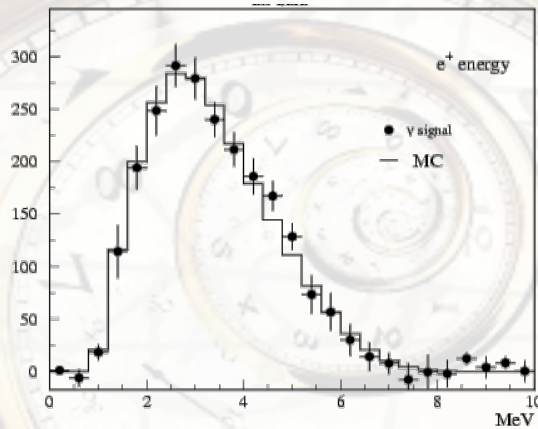


# Reactor Neutrinos

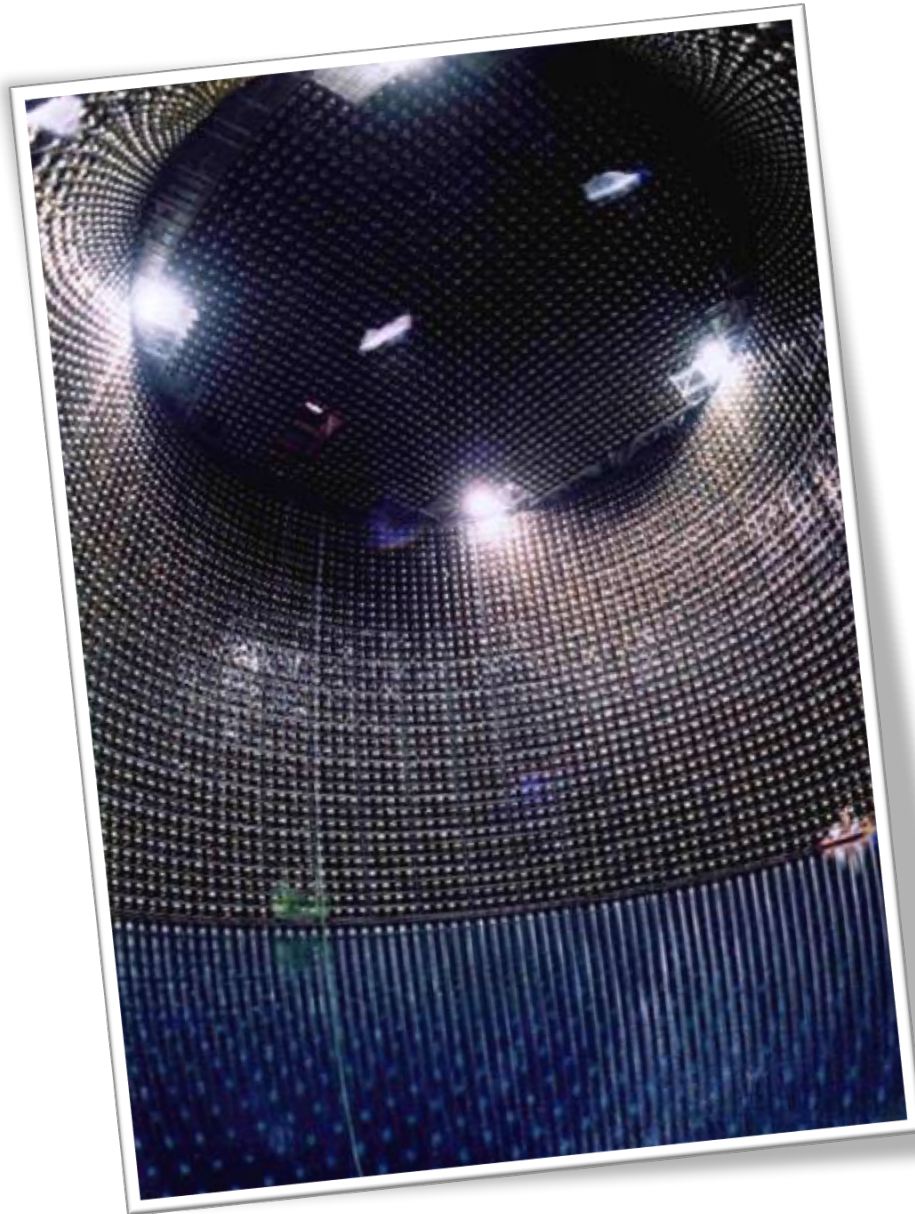
## 1995 - 2018



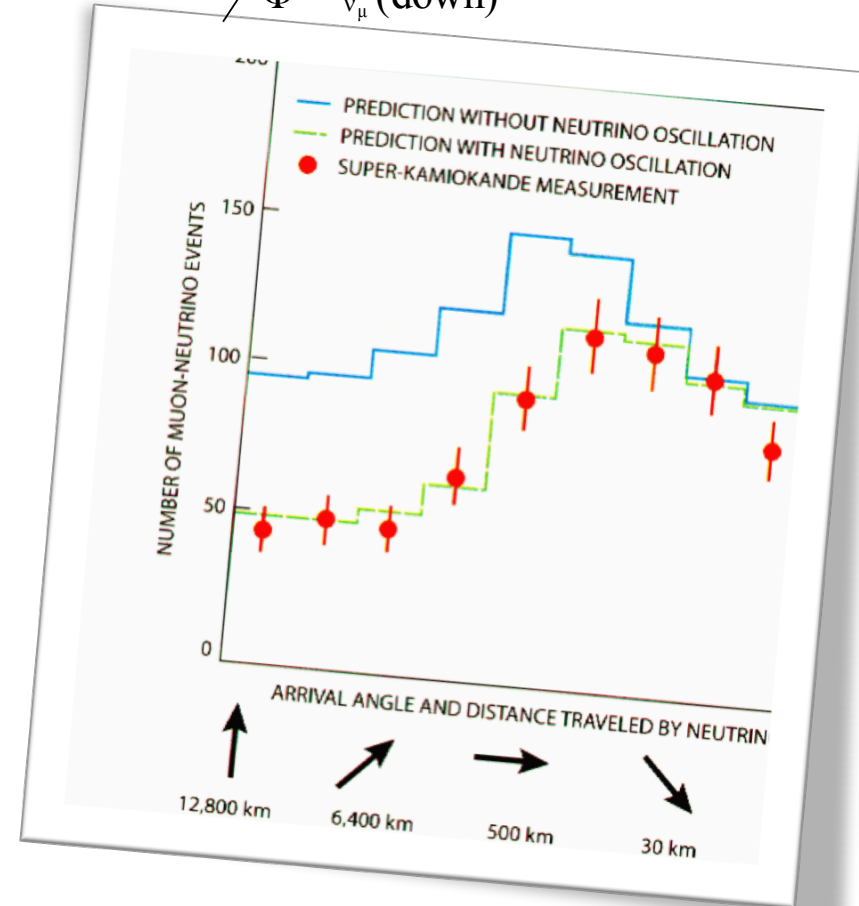
International conference on the History of the Neutrino  
Paris, 07/09/2018

Thierry Lasserre  
CEA-Saclay – APC – TUM/MPP

# Context: Atmospheric Neutrino Anomaly

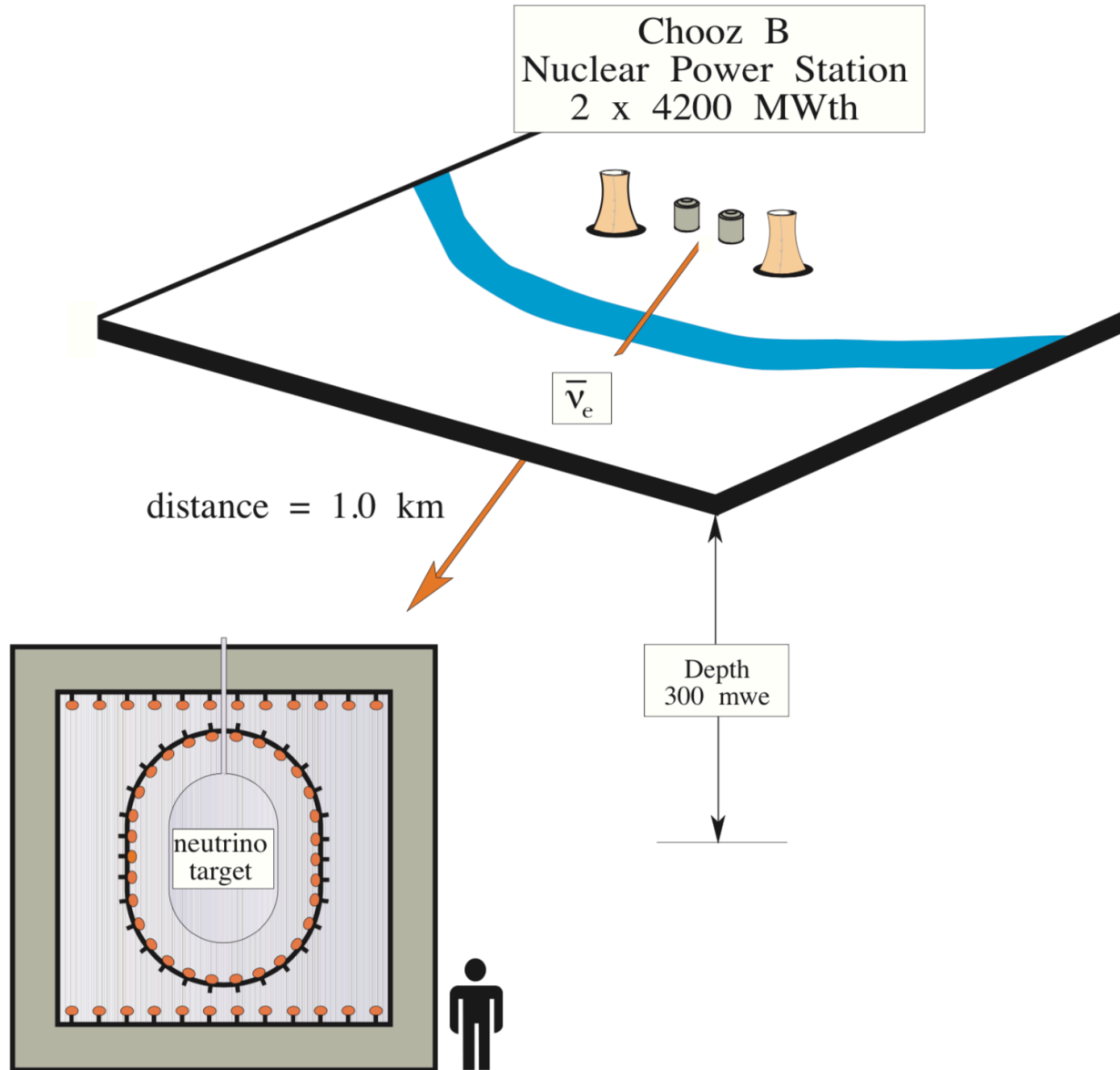


$$\frac{\Phi^{\text{Atm}}_{\nu_{\mu}}(\text{up})}{\Phi^{\text{Atm}}_{\nu_{\mu}}(\text{down})} = 0.54 \pm 0.04$$



Neutrino do have mass and mix

# CHOOZ (1997-1998)



- CHOOZ reactor, Ardennes (France)  
Underground gallery (300 mwe) found by the Mother of H. de Kerret!



- First large underground monolithic detector for reactor neutrinos.
- 2 cores: 2 x 4.2 GW
- 5 tons of Gd-doped liquid scintillator (degradation of its attenuation over time)

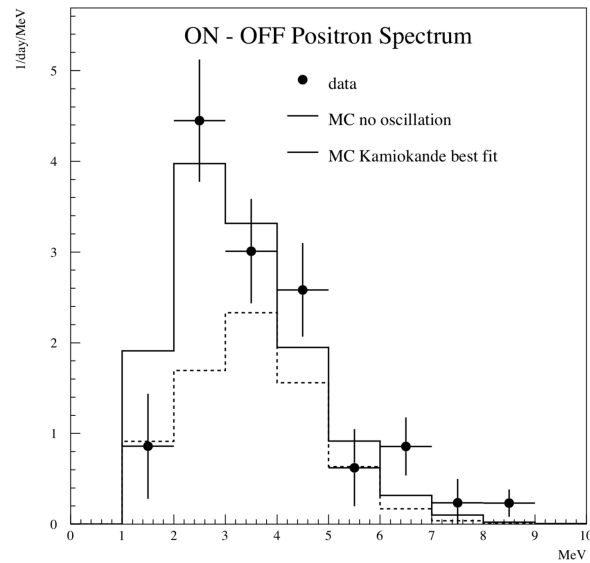




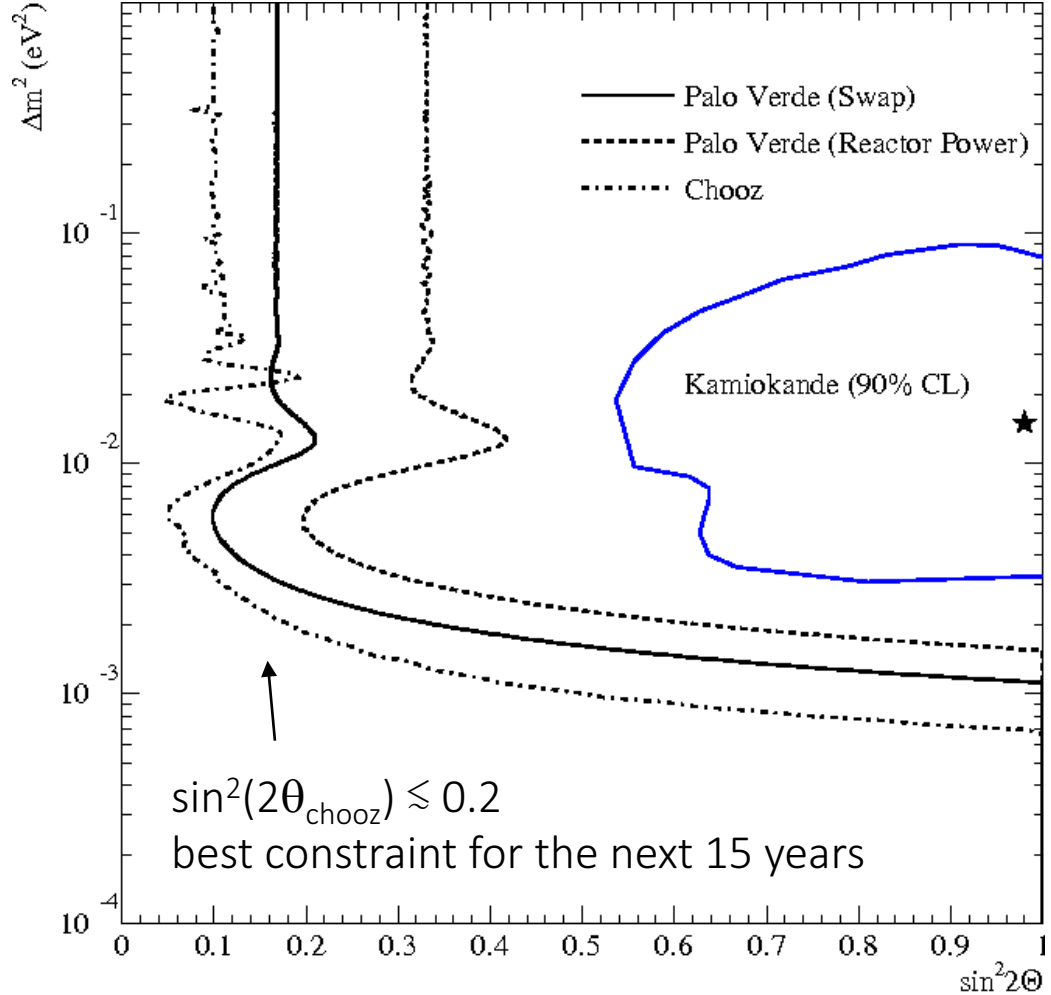
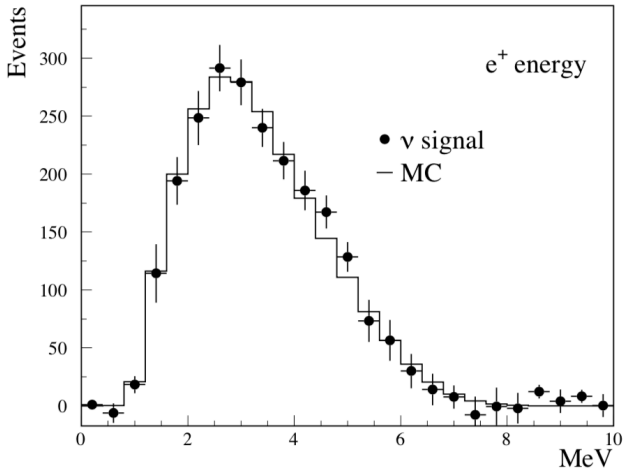
# Results

Neither Chooz nor Palo-Verde reported oscillation. This null-result eliminated  $\nu_\mu \rightarrow \nu_e$  as the primary mechanism for the (Super-)Kamiokande atmospheric deficit

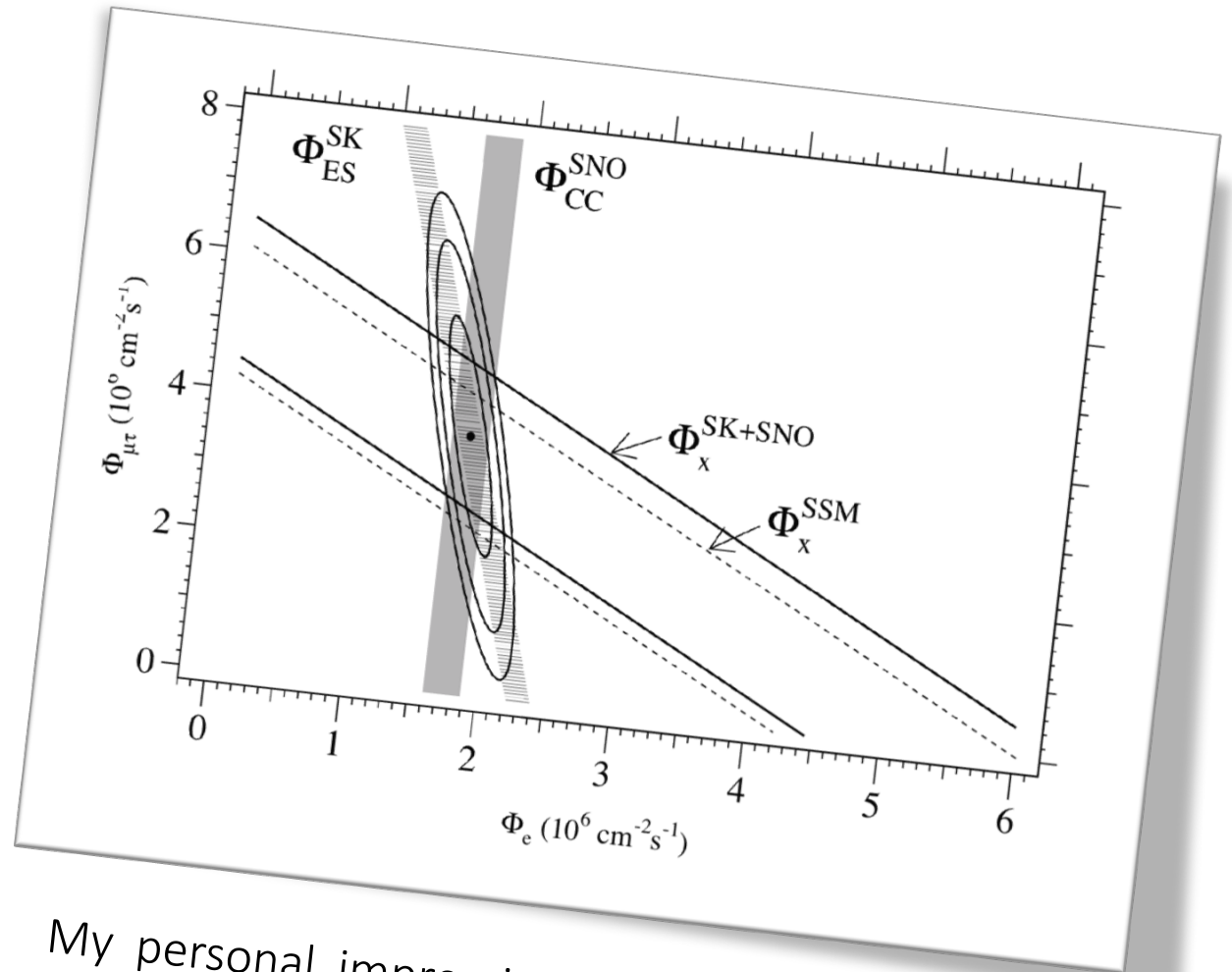
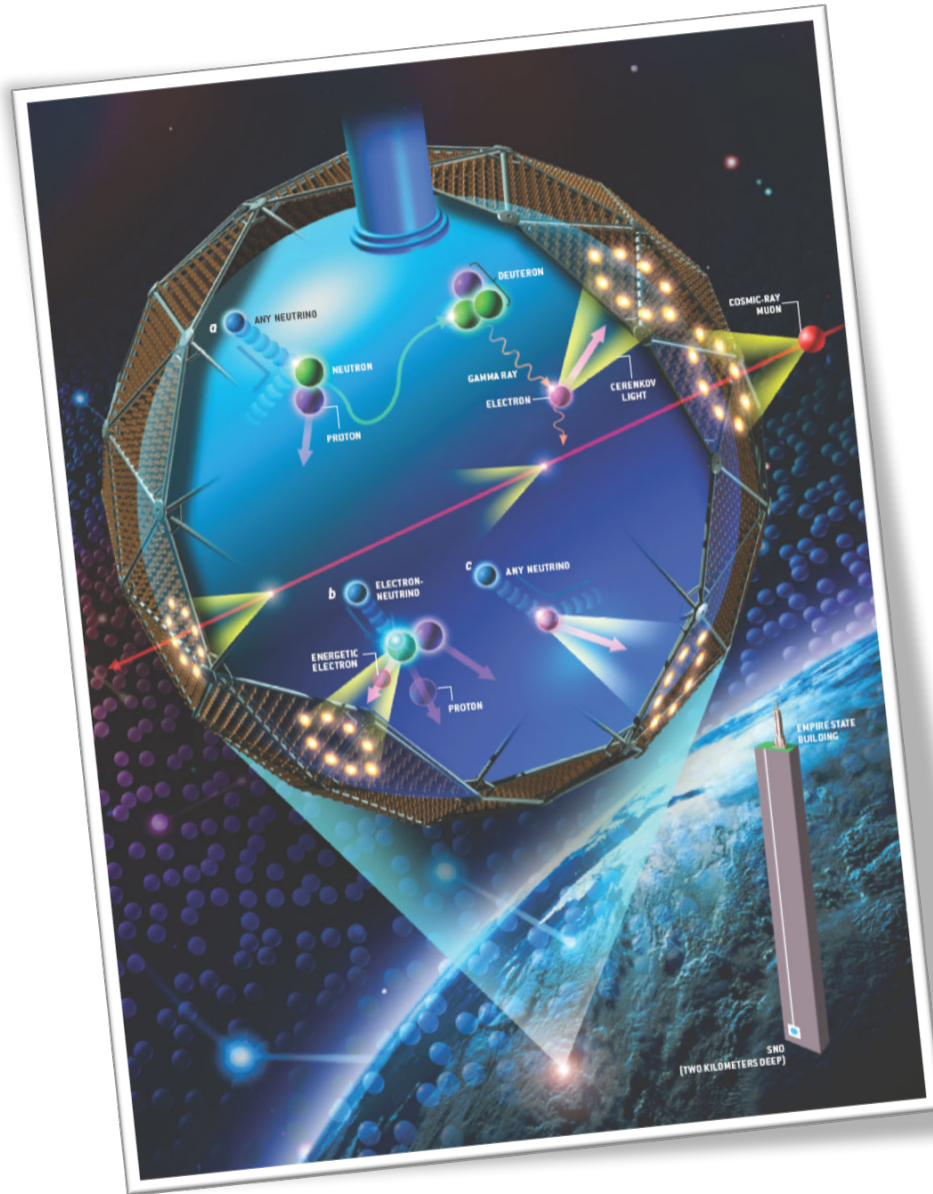
Palo-Verde  
1000 evts



Chooz  
1000 evts

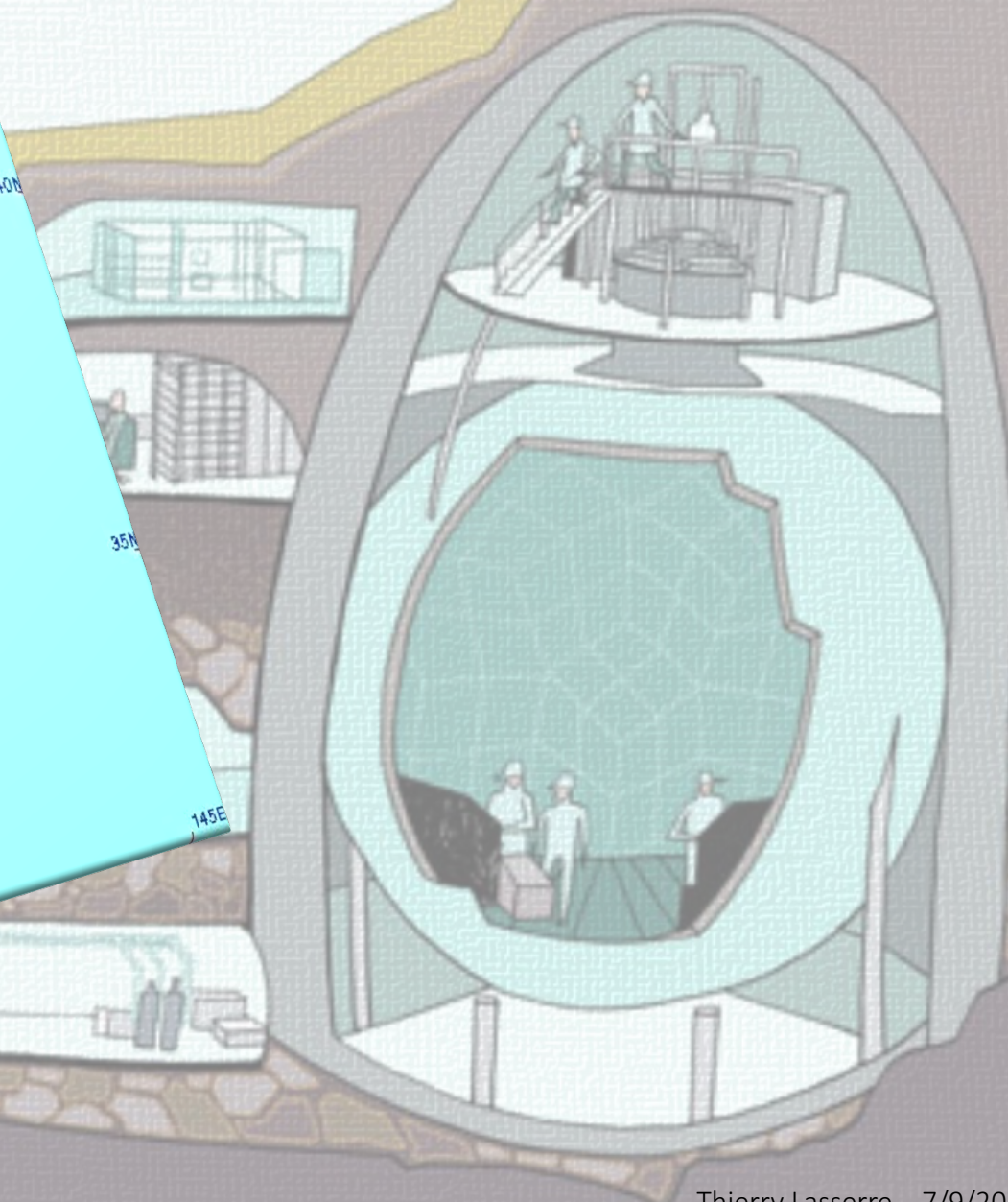
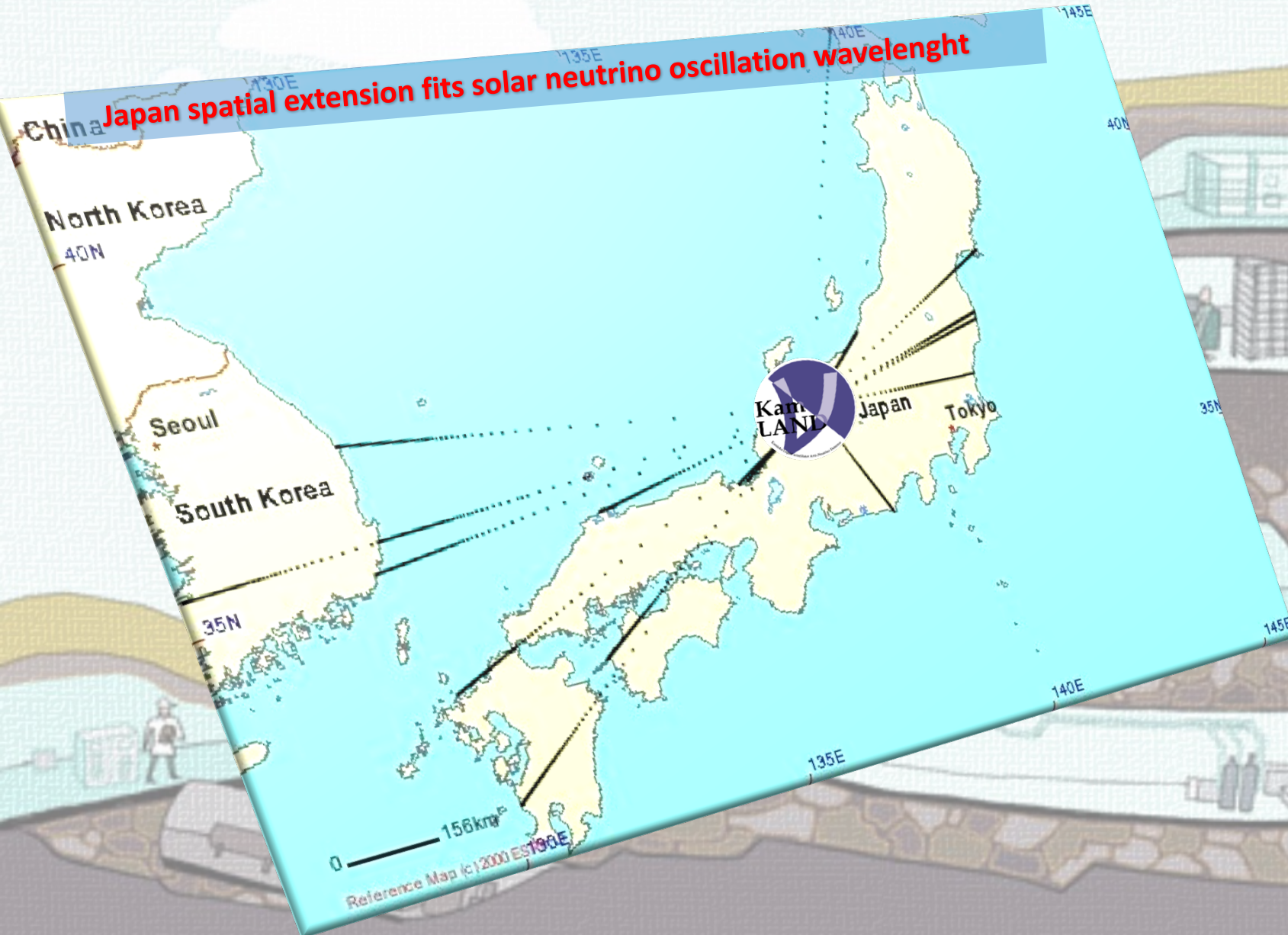


# Context: Solar neutrinos till first SNO results (2001)



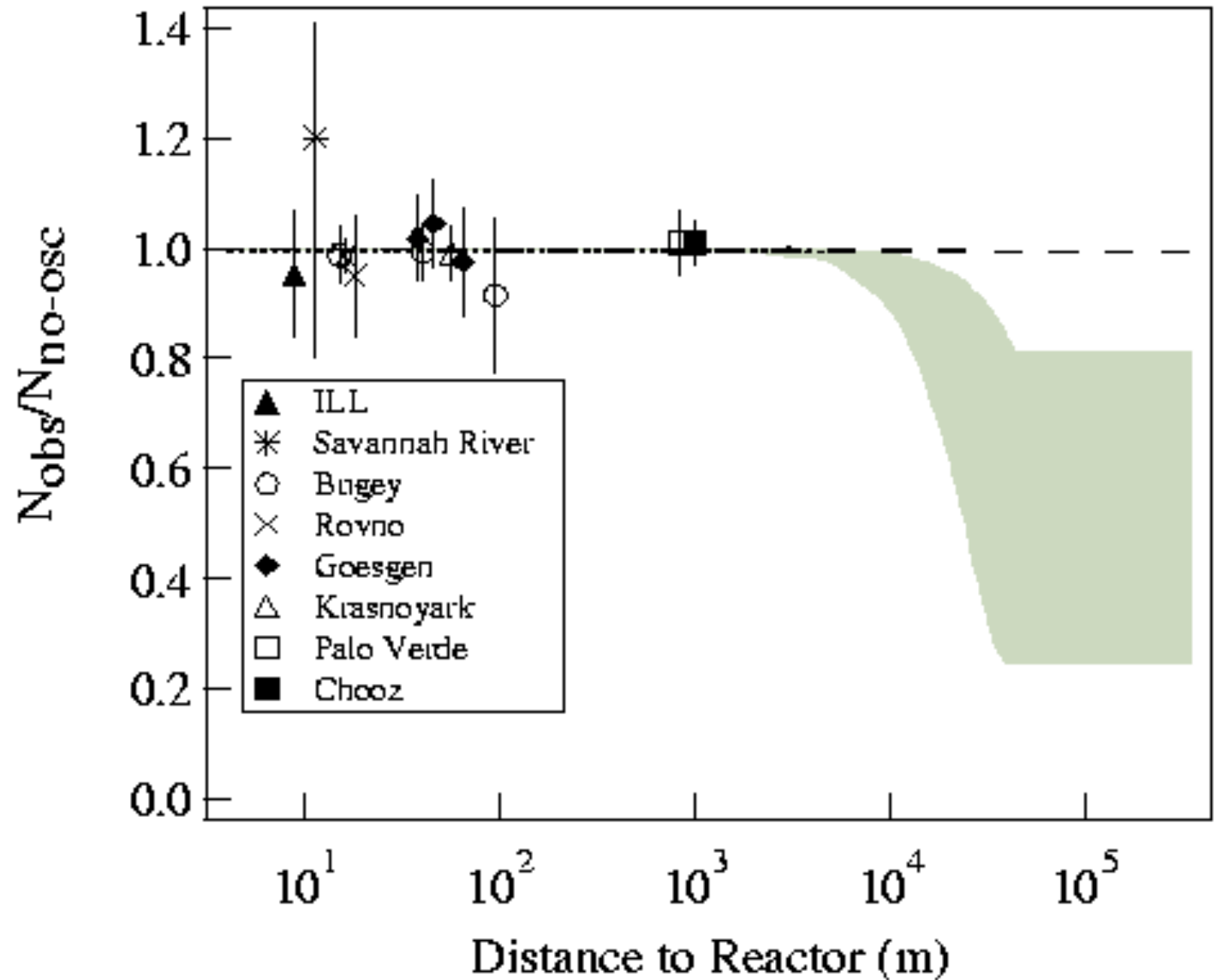
My personal impression at that time: a sizable fraction of the community considered neutrino beams as the –only– future of neutrino physics

# Japan and Neutrino Oscillations: A perfect match...

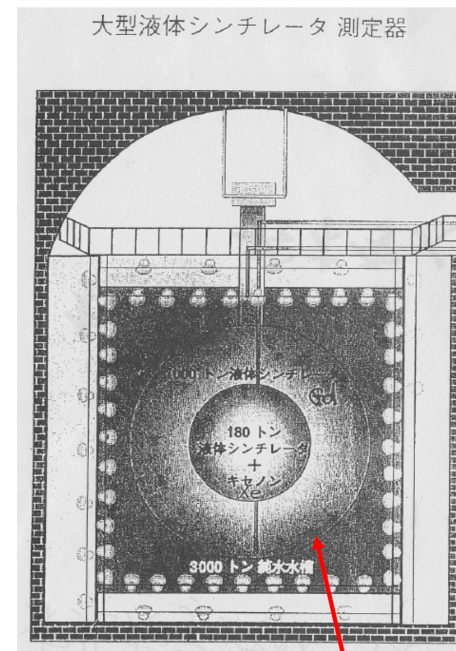
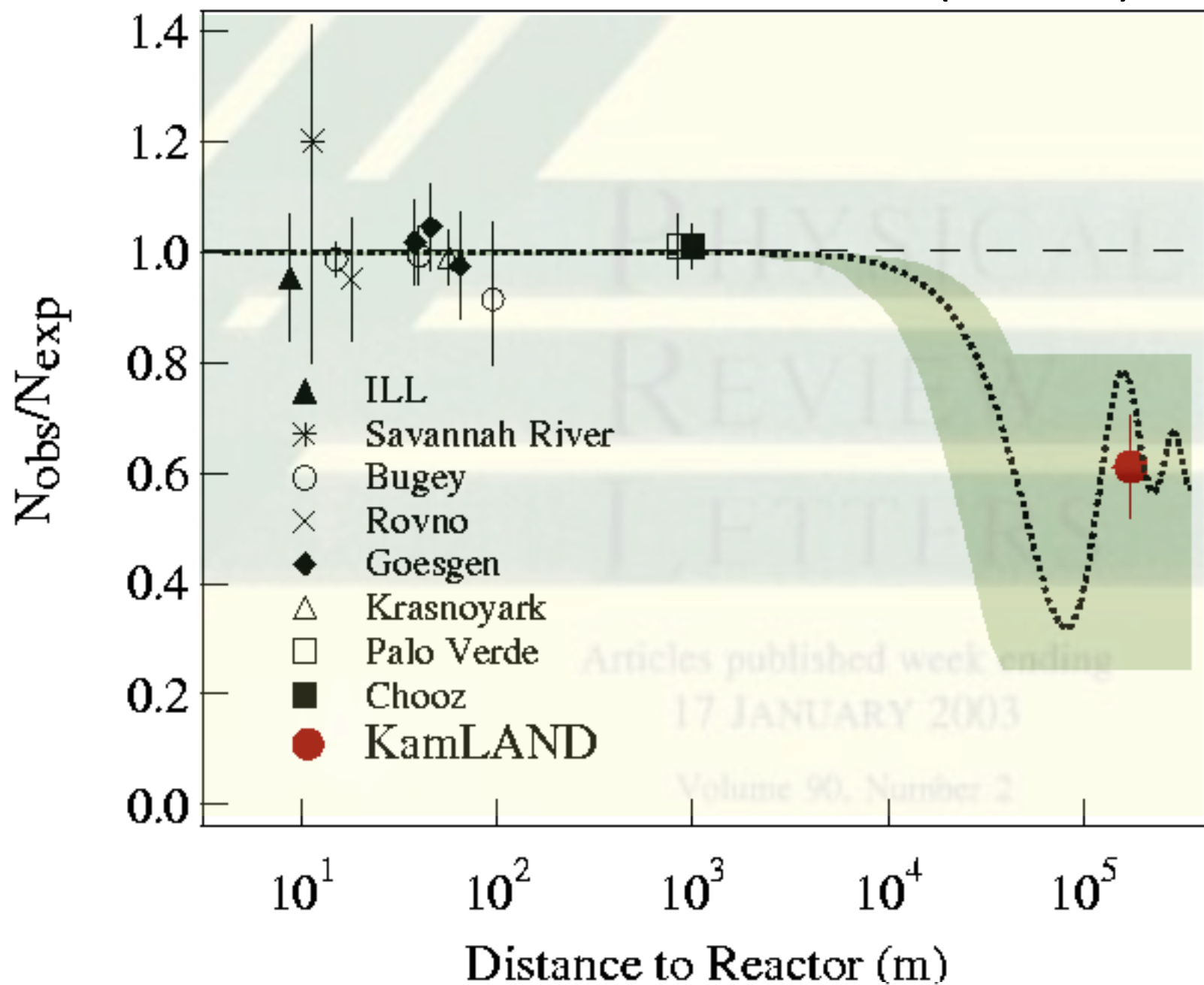




# First KamLAND Results (2002)

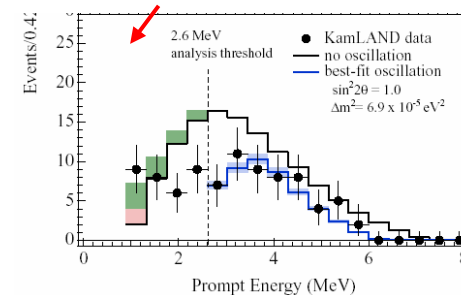


# First KamLAND Results (2002)

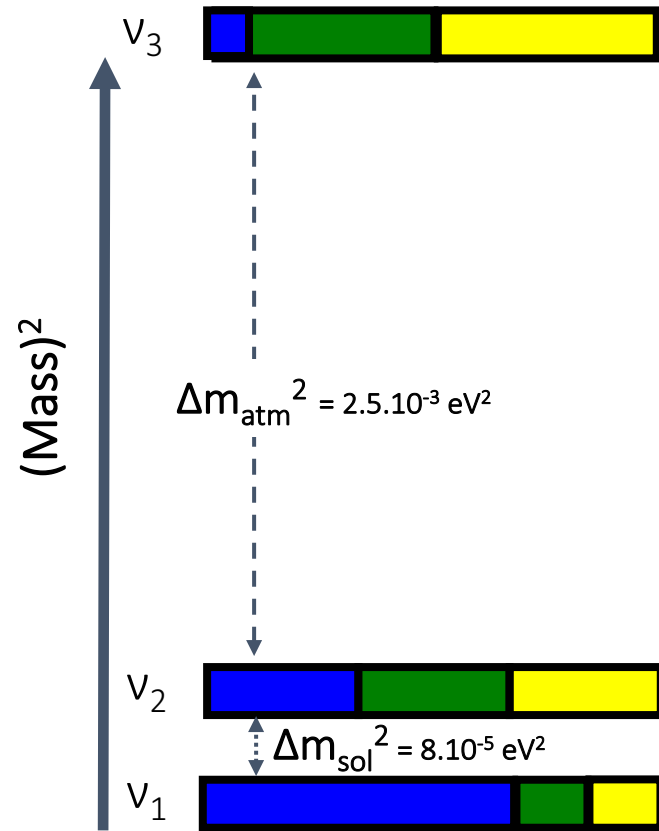


1995 design

2002 results



# End of 2002



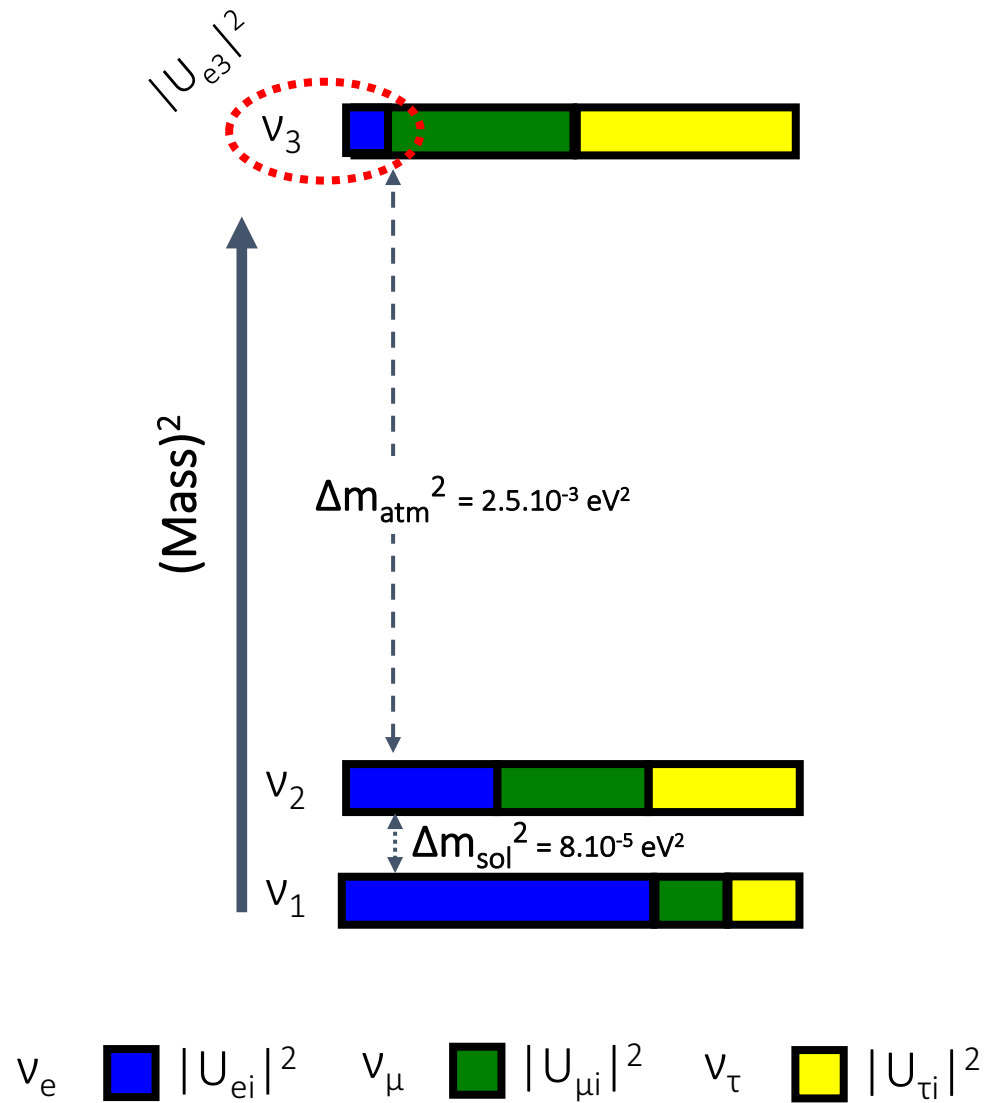
$$\sin(2\theta_{13})^2 < 0.2$$

$$\sin(2\theta_{23})^2 \geq 0.9$$

$$\sin(2\theta_{12})^2 \approx 0.8$$

$\nu_e$    $|U_{ei}|^2$     $\nu_\mu$    $|U_{\mu i}|^2$     $\nu_\tau$    $|U_{\tau i}|^2$

# $\theta_{13}$ made it to the top priority list

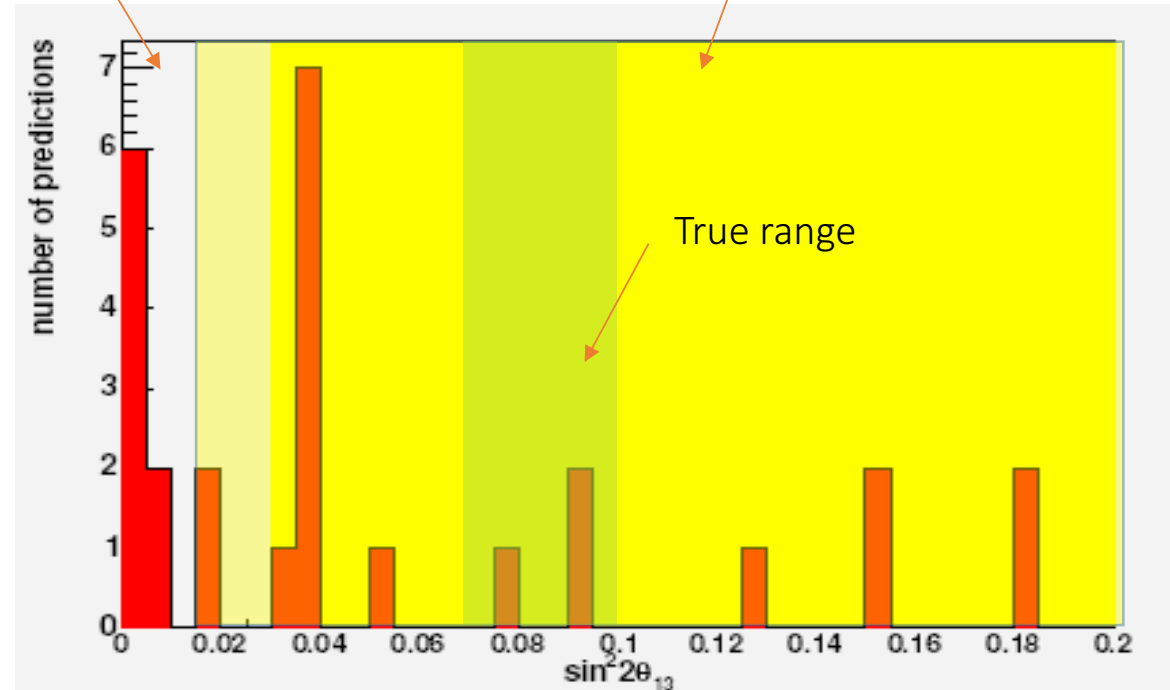


# By 2002...

Reference	$\sin^2\theta_{13}$	if $\sin^2 2\theta_{13} < 0.01$
<i>SO(10)</i>		
Goh, Mohapatra, Ng [40]	0.18	<del>0.18</del>
<i>Orbifold SO(10)</i>		
Asaka, Buchmüller, Covi [41]	0.1	<del>0.04</del>
<i>SO(10) + flavor symmetry</i>		
Babu, Pati, Wilczek [42]	$5.5 \cdot 10^{-4}$	$1.2 \cdot 10^{-6}$
Blazek, Raby, Tobe [43]	0.05	0.01
Kitano, Mimura [44]	0.22	<del>0.18</del>
Albright, Barr [45]	0.014	$7.8 \cdot 10^{-4}$
Maekawa [46]	0.22	<del>0.18</del>
Ross, Velasco-Sevilla [47]	0.07	<del>0.02</del>
Chen, Mahanthappa [48]	0.15	<del>0.09</del>
Raby [49]	0.1	<del>0.04</del>
<i>SO(10) + texture</i>		
Buchmüller, Wyler [50]	0.1	<del>0.04</del>
Bando, Obara [51]	0.01 .. 0.06	$4 \cdot 10^{-4}$ .. 0.01
<i>Flavor symmetries</i>		
Grimus, Lavoura [52, 53]	0	0
Grimus, Lavoura [52]	0.3	<del>0.3</del>
Babu, Ma, Valle [54]	0.14	<del>0.08</del>
Kuchimanchi, Mohapatra [55]	0.08 .. 0.4	<del>0.03 .. 0.3</del>
Ohlsson, Seidl [56]	0.07 .. 0.14	<del>0.02 .. 0.08</del>
King, Ross [57]	0.2	<del>0.15</del>
<i>Textures</i>		
Honda, Kaneko, Tanimoto [58]	0.08 .. 0.20	<del>0.03 .. 0.15</del>
Lebed, Martin [59]	0.1	<del>0.04</del>
Bando, Kaneko, Obara, Tanimoto [60]	0.01 .. 0.05	$4 \cdot 10^{-4}$ .. 0.01
Ibarra, Ross [61]	0.2	<del>0.15</del>
<i>3 x 2 see-saw</i>		
Appelquist, Piai, Shrock [62, 63]	0.05	0.01
Frampton, Glashow, Yanagida [64]	0.1	<del>0.04</del>
Mei, Xing [65] (normal hierarchy)	0.07	<del>0.02</del>
(inverted hierarchy)	> 0.006	> $1.6 \cdot 10^{-4}$
<i>Anarchy</i>		
de Gouvêa, Murayama [66]	> 0.1	<del>&gt; 0.04</del>
<i>Renormalization group enhancement</i>		
Mohapatra, Parida, Rajasekaran [67]	0.08 .. 0.1	<del>0.03 .. 0.04</del>

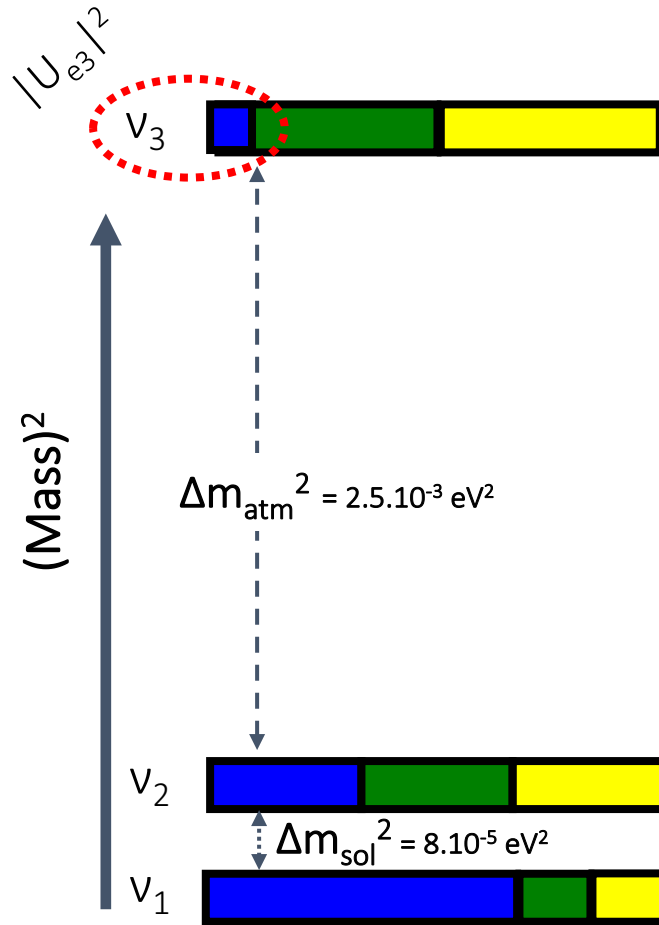
Need for  
v-factories

Sensitivity goal for reactor experiments



The difference between theory and practice is larger in practice than the difference between theory and practice in theory.

# $\theta_{13}$ made it to the top priority list



How to measure  $\theta_{13}$  ?

connect  $\nu_e$  state to the 'isolated' neutrino  $\Delta m_{\text{atm}}^2$

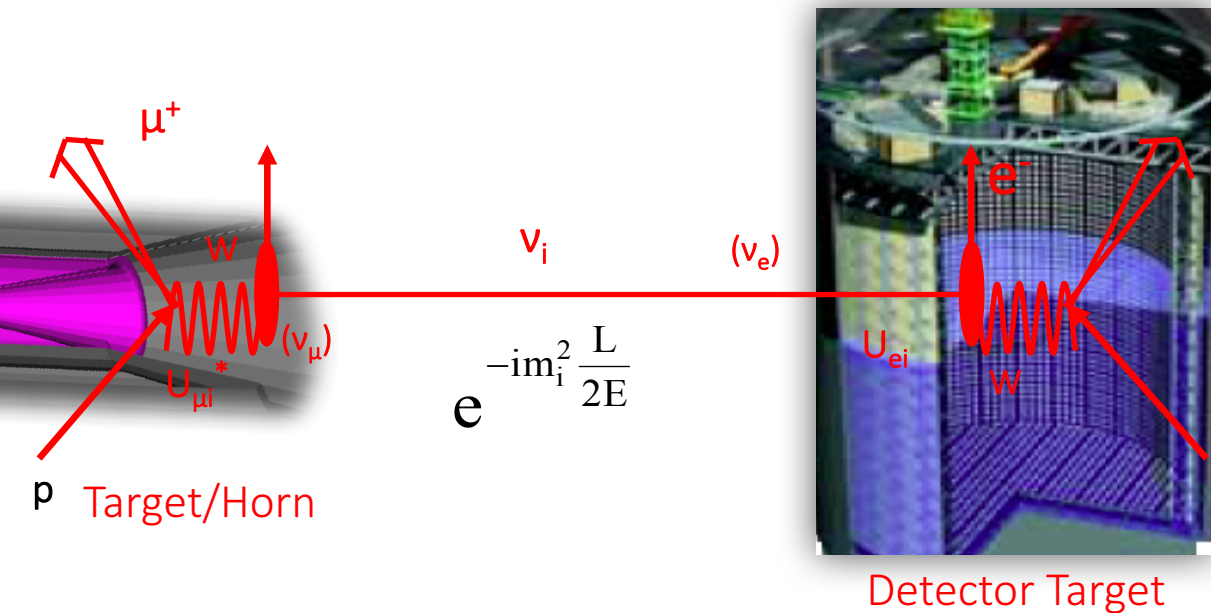
- $L = 1\text{-}2 \text{ km}$ ,  $E = \text{MeV}$ 
  - disappearance channel. @reactor
  - $\theta_{13}$  only

or

- $L = 100\text{-}1000 \text{ km}$ ,  $E = \text{GeV}$ 
  - appearance channel
  - $\theta_{13}$ , NH/IH,  $\delta_{\text{CP}}$  are entangled

$\nu_e$    $|U_{ei}|^2$      $\nu_\mu$    $|U_{\mu i}|^2$      $\nu_\tau$    $|U_{\tau i}|^2$

# Beam

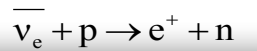
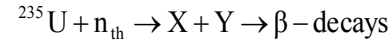
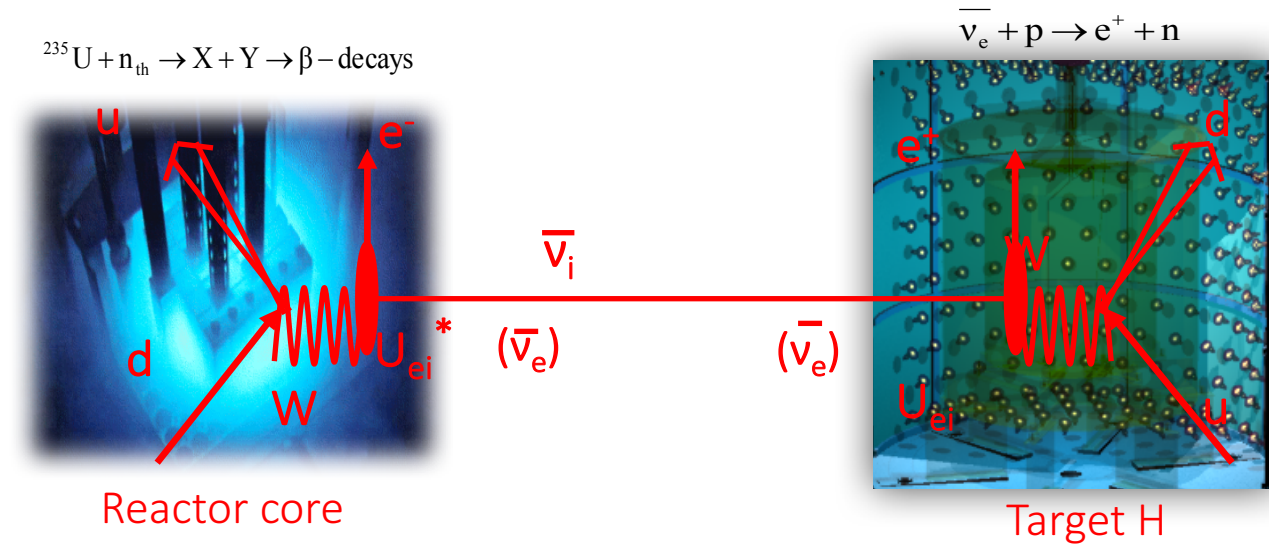


$$e^{-im_i^2 \frac{L}{2E}} \nu_i \rightarrow \nu_e$$

$$P(\nu_\mu \rightarrow \nu_e) = \left[ \sum_i U_{\mu i}^* e^{-im_i^2 \frac{L}{2E}} U_{ei} \right]^2$$

- Complex oscillation formula
  - depends on  $\sin^2(2\theta_{13})$ ,  $\Delta m_{31}^2$ ,  $\text{sign}(\Delta m_{31}^2)$ ,  $\delta$
- appearance experiments
  - $\sin^2(2\theta_{13})$  measurement depends on  $\delta$ -CP
- matter effects
  - High potential for detecting  $\theta_{13}$ , less for measuring it precisely

# Reactor



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2(2\theta_{13}) \left[ \sin^2 \left( 1.27 \frac{\Delta m_{\text{atm}}^2 (\text{eV}^2) L (\text{m})}{E (\text{MeV})} \right) + \mathcal{O} \left( \frac{\Delta m_{\text{sol}}^2}{\Delta m_{\text{atm}}^2} \right) \right]$$

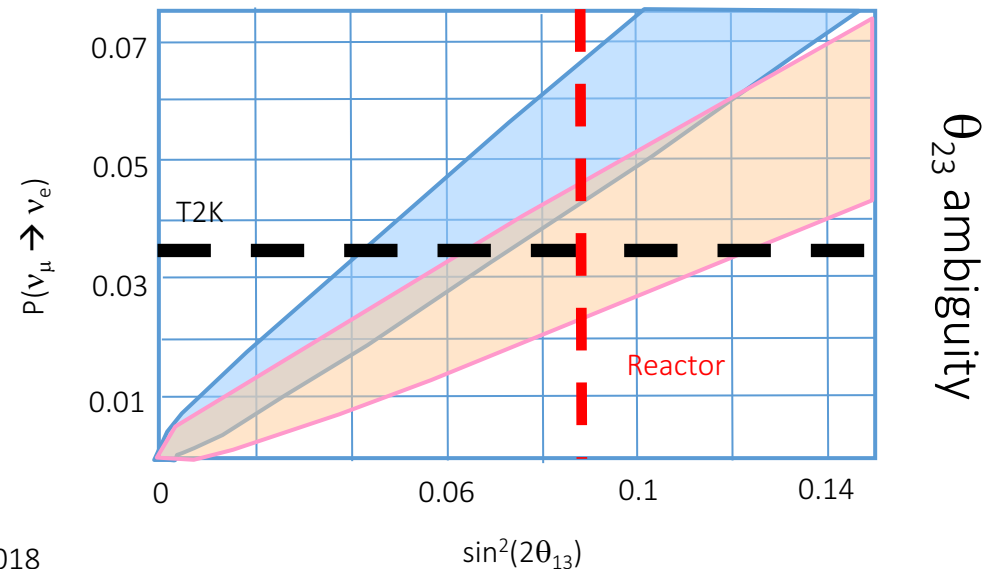
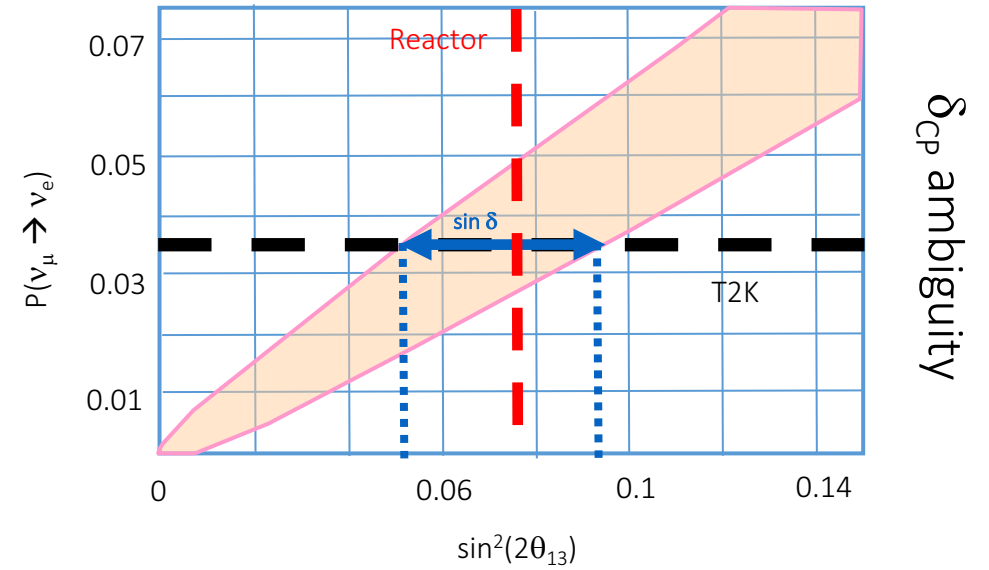
- Simple oscillation formula
  - depends  $\sin^2(2\theta_{13})$  &  $\Delta m_{\text{atm}}^2$ , weakly on  $\Delta m_{\text{sol}}^2$
- disappearance experiment
  - $\sin^2(2\theta_{13})$  measurement independent of  $\delta$ -CP
- no matter effects
  - $\sin^2(2\theta_{13})$  measurement independent of mass hierarchy

# Some Sensitivity predictions (2002)

Table laid down on the CEA-DAPNIA director desk on Oct. 2002

Experiment	$\sin^2(2\theta_{13})$ 90% C.L.	Expected start
CHOOZ	<0.14	-
MINOS	<0.06	?
ICARUS	<0.04	2007 ?
OPERA	<0.06	2007 ?
JHF2K (T2K)	<0.006	2009 ?
Kr2Det - Russia	<0.02	?
Kaska - Japan	>0.02	?
XXX – France/EU With a near detector	< 0.025	2006 ?

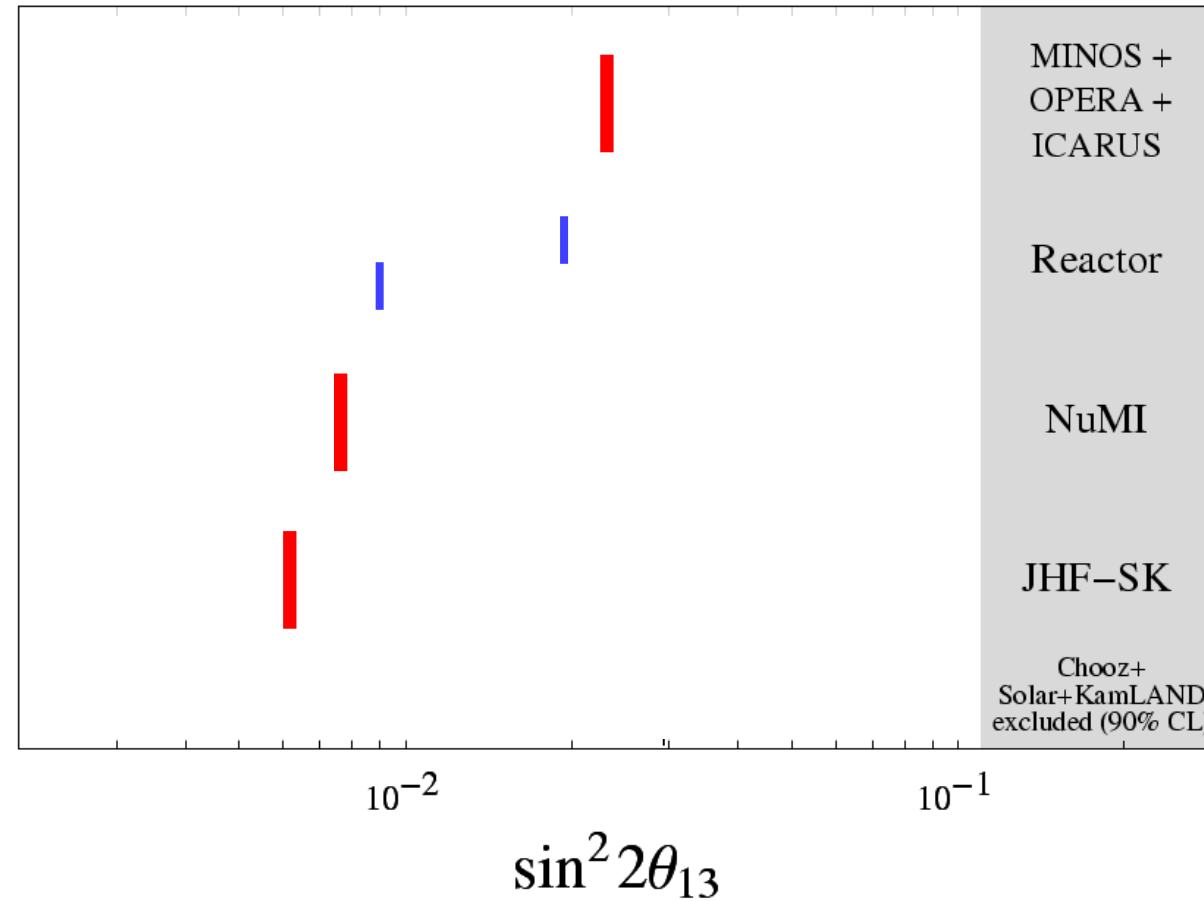
$$\Delta m^2_{\text{atm}} = 2.5 \cdot 10^{-3} \text{ eV}^2$$





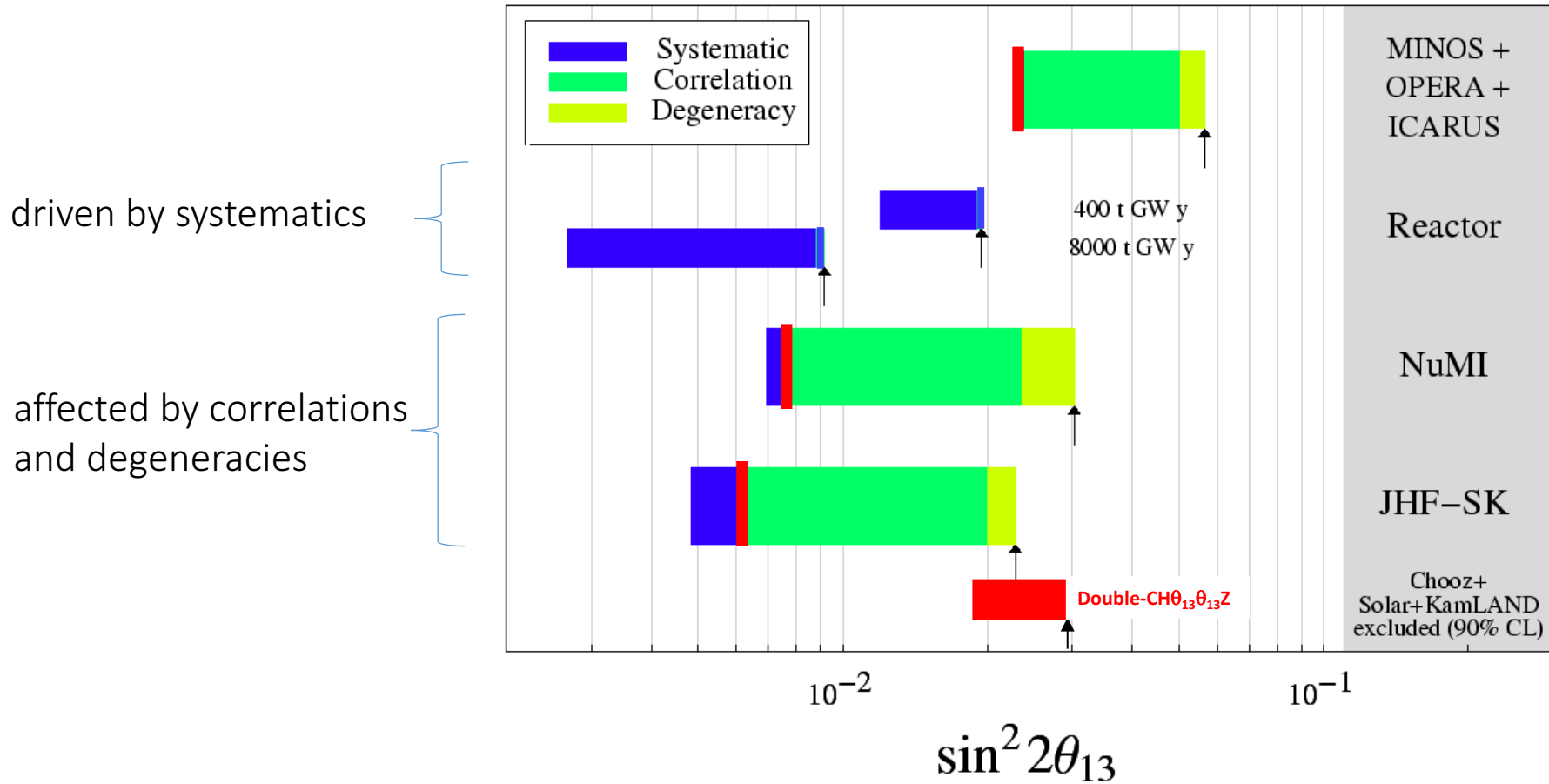
# Reactor Versus Beam Experiments

Sensitivity to  $\sin^2 2\theta_{13}$  at 90% CL



# Towards a fair comparison

Sensitivity to  $\sin^2 2\theta_{13}$  at 90% CL

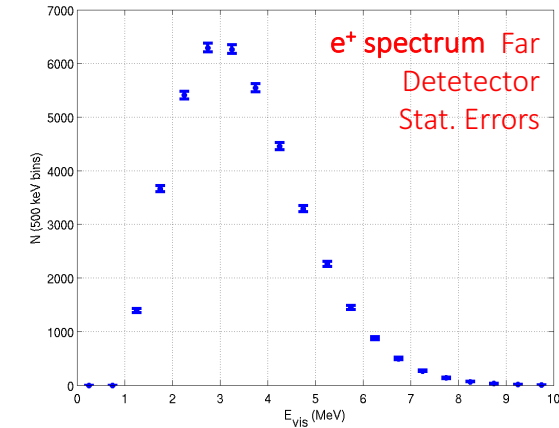
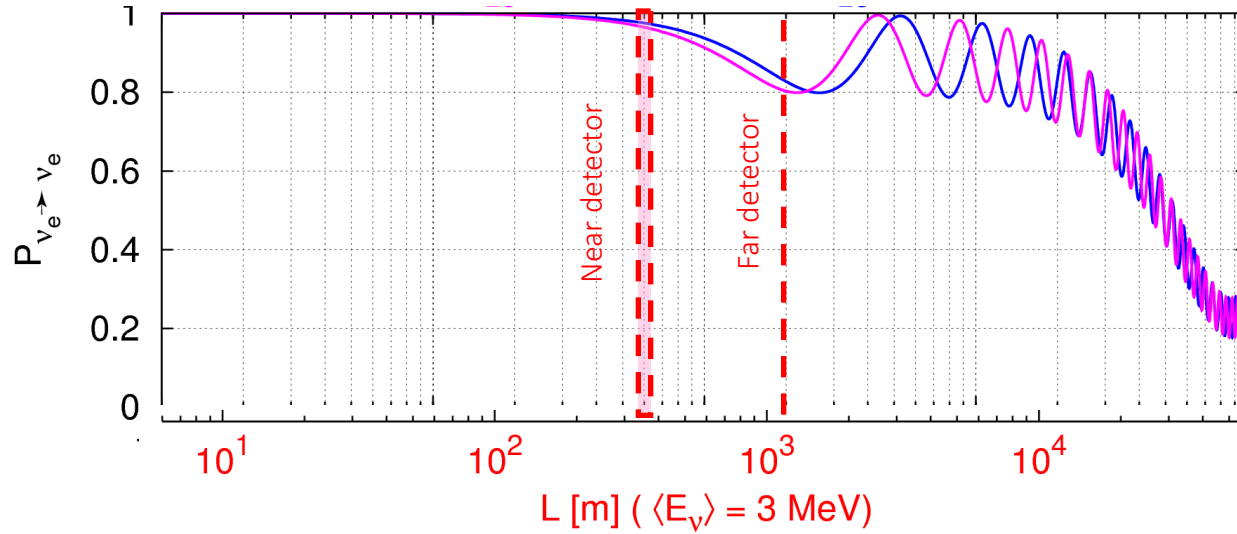


Nuclear Physics B 665 (2003) 487–519

CP-violation asymmetry 'proportional' to  $\sin \theta_{13}$ . Synergy was finally established and accepted

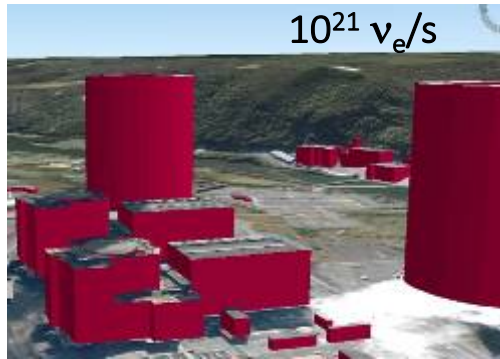
# The - new - concept – no so new, indeed!

$$P(\nu_e \rightarrow \nu_e) = 1 - \sin^2(2\theta_{13}) \sin^2(1.27 \Delta m_{31}^2 L/E)$$

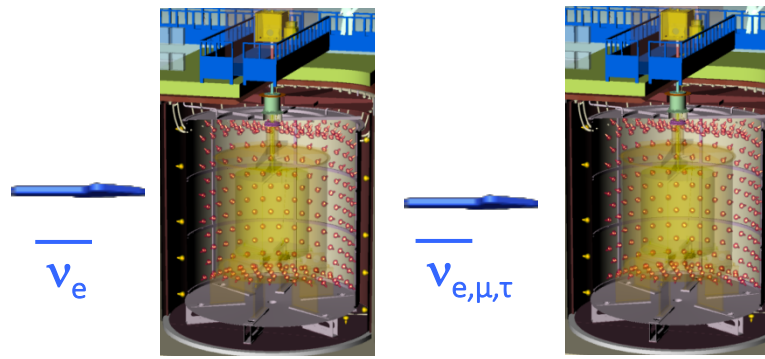


$$\Delta m_{\text{atm}}^2 = 3.0 \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta_{13}) = 0.12$$

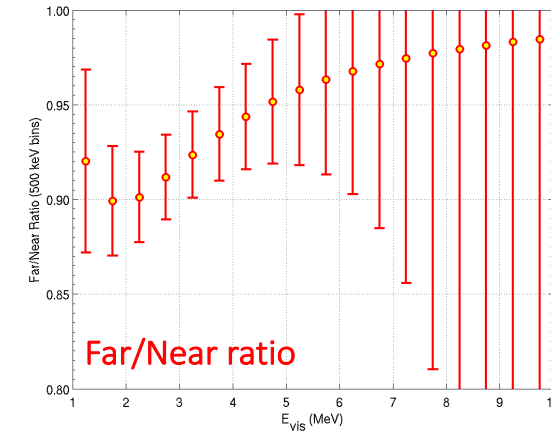


Nuclear Power Station



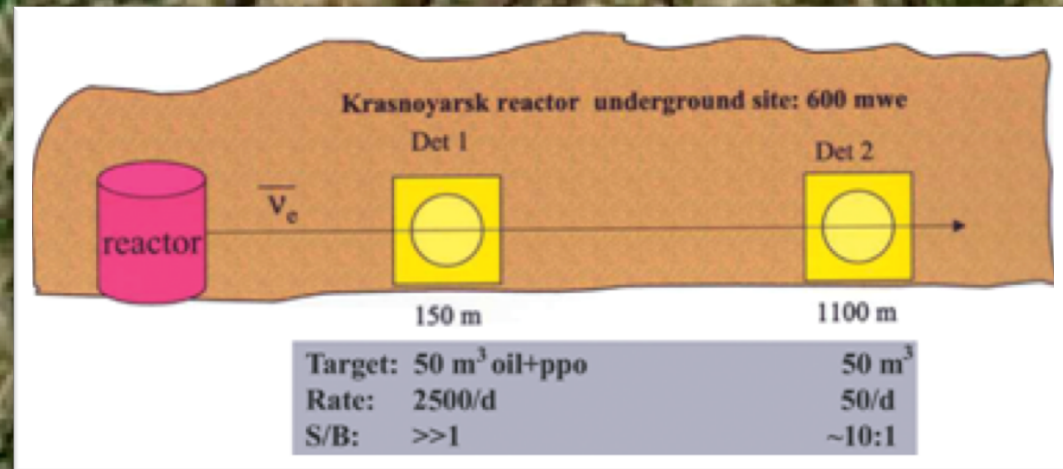
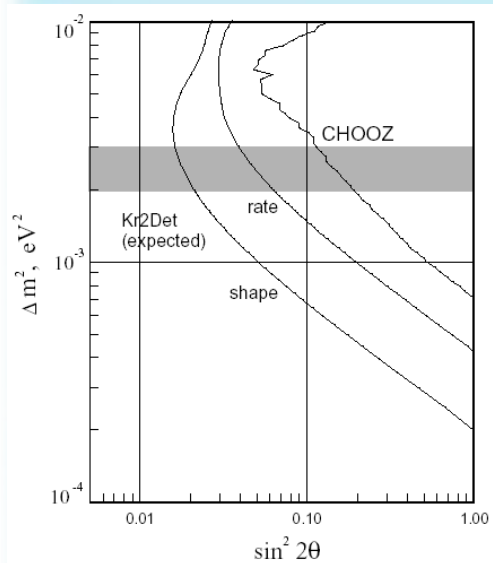
Near  
detector  
<500 m

Far  
detector  
1-2 km



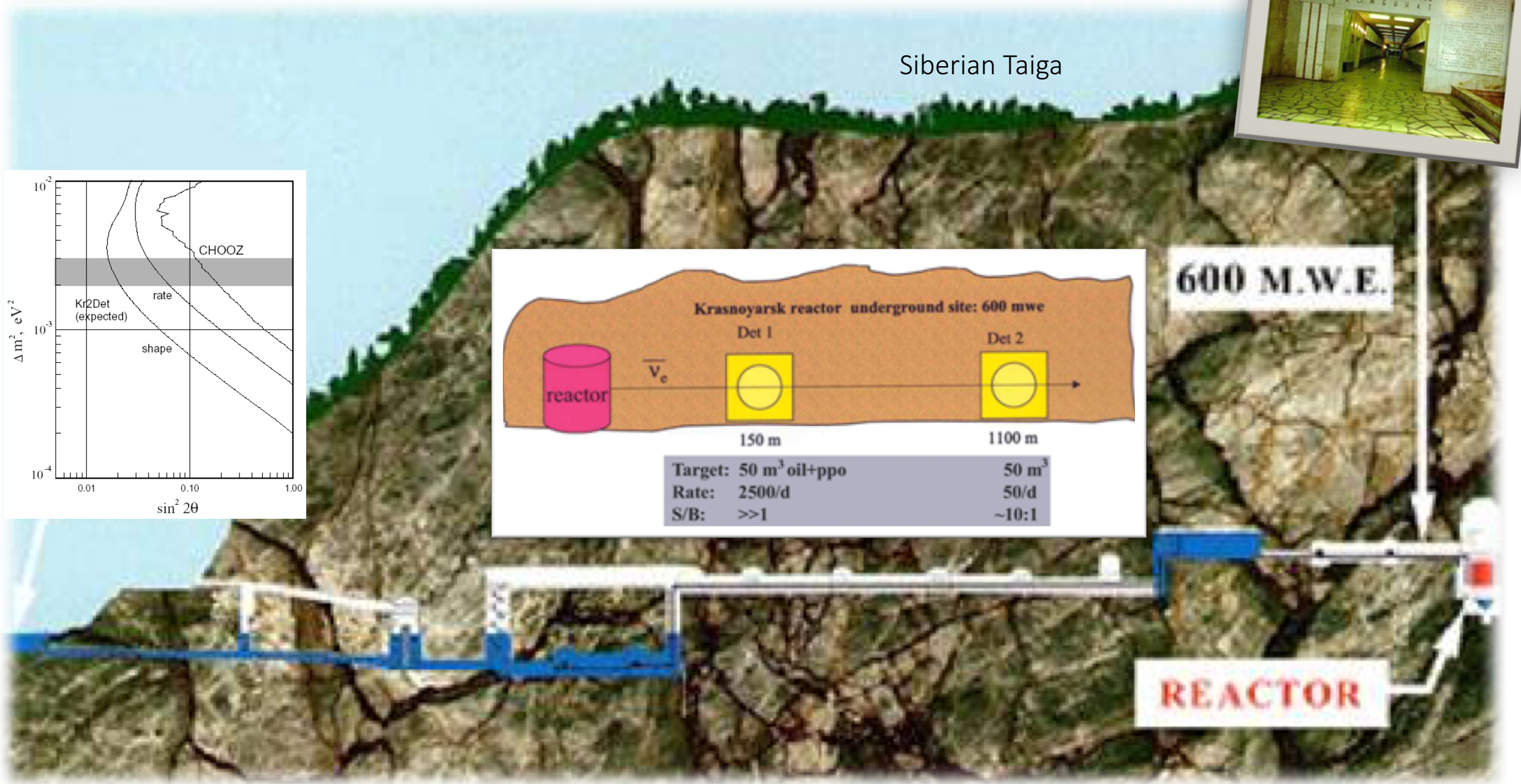
# Kr2Det (Krasnoyarsk, Russia)

Siberian Taiga



600 M.W.E.

REACTOR



# Workshop on Future Low-Energy Neutrino Experiments

April 30-May 2, 2003

Bryant Conference Center

University of Alabama

Tuscaloosa, Alabama

Sponsored by the University of Alabama and Argonne National Laborator

## Purpose

The workshop will focus on the study of neutrino oscillations by future reactor experiments. It will be attended by participants in the Kr2Det and Kashiwasaki proposals, and a major part of the workshop will be devoted to the presentation and discussion of these proposals. Presentations which describe other proposals or a discussion of these proposals.



## Workshop on Future Low-Energy Neutrino Experiments

October 9-11, 2003

Workshop on Future Low-Energy Neutrino Experiments, Munich, October 9-11, 2003

### Committees

Program Committee	Local Organizing Committee
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Stefan Schoenert (MPK, Heidelberg)	
Fumihiko Suekane (Tohoku University, Sendai)	

- 12/2002: 1<sup>st</sup> EU meeting, MPIK Heidelberg
- 04/2003: 2<sup>nd</sup> EU meeting, PCC, Paris
- 05/2003: 1<sup>st</sup> international meeting, Alabama Univ, USA
- 10/2003: 2<sup>nd</sup> international meeting, TUM, Germany
- 03/2004: 3<sup>rd</sup> international meeting, Niigata, Japan (first presentation of Daya Bay)

... Put 3 experiments on-track: Daya Bay, Double Chooz, RENO



## 3rd Workshop on Future Low-Energy Neutrino Experiments

Date: March 20(Sat)-22(Mon), 2004

Place: Toki-Messe, Niigata, Japan

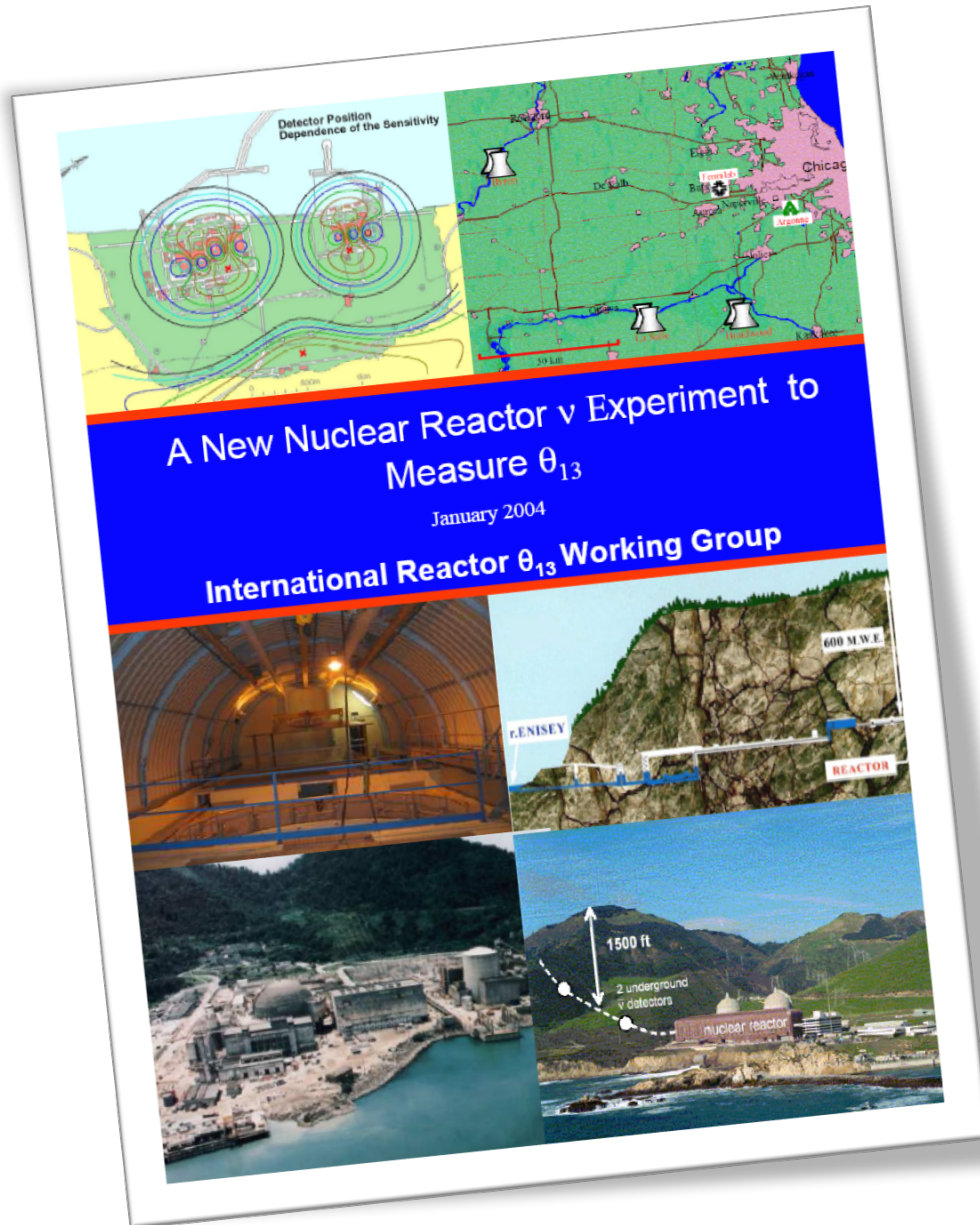
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<http://www.hep.sc.niigata-u.ac.jp/~neutrino/workshop/>

# $\theta_{13}$ White Paper - hep-ex/0402041



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 R. Svoboda<sup>20</sup> R. Talaga<sup>3</sup> N. Tamura<sup>25</sup>  
 M. Tanimoto<sup>25</sup> J. Thron<sup>3</sup> E. von Toerne<sup>18</sup>  
 D. Vignaud<sup>15</sup> C. Wagner<sup>3</sup> Y.F. Wang<sup>17</sup> Z. Wang<sup>17</sup>  
 W. Winter<sup>33</sup> H. Wong<sup>1</sup> E. Yakushev<sup>2</sup> C.G. Yang<sup>17</sup>  
 O. Yasuda<sup>37</sup>

125 authors  
40 institutions  
9 countries

January 26, 2004

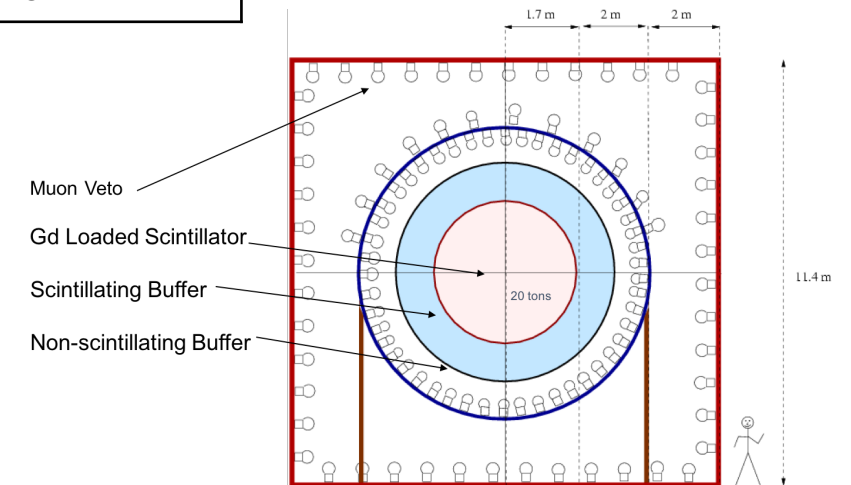
# Challenges towards $\sin^2(2\theta_{13}) \approx 0.01$

Best Sensitivity (Chooz): obs/pred =  $1.01 \pm 2.8\%$ (stat)  $\pm 2.7\%$ (syst)

Statistics	CHOOZ	DoubleChooz	DayaBay
Baseline	1 km	1 km	2 km
Reactor Power	2 x 8 GW	x 1	X 4
Detector Volume	5 m <sup>3</sup>	X 2	x 16
Data Taking Time	0.5 y	X 10	X 10
Candidates	10 <sup>3</sup>	10 <sup>5</sup>	10 <sup>6</sup>

Baseline optimization

Systematics	CHOOZ	DoubleChooz
Reactor Flux	1.9 %	---
Detection Efficiency	0.8 %	< 0.2 %
Energy scale		<< 1%
Target Volume	1.5 %	< 0.5 %



**Background reduction: accidentals / 100...**

New 4-region large detector concept  
Proposed by Double Chooz (2003)

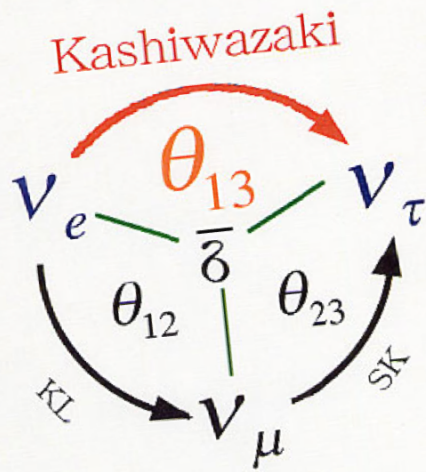


A plethora of sites investigated (2002-2005)



# KASA (Kashiwasaki, Japan)

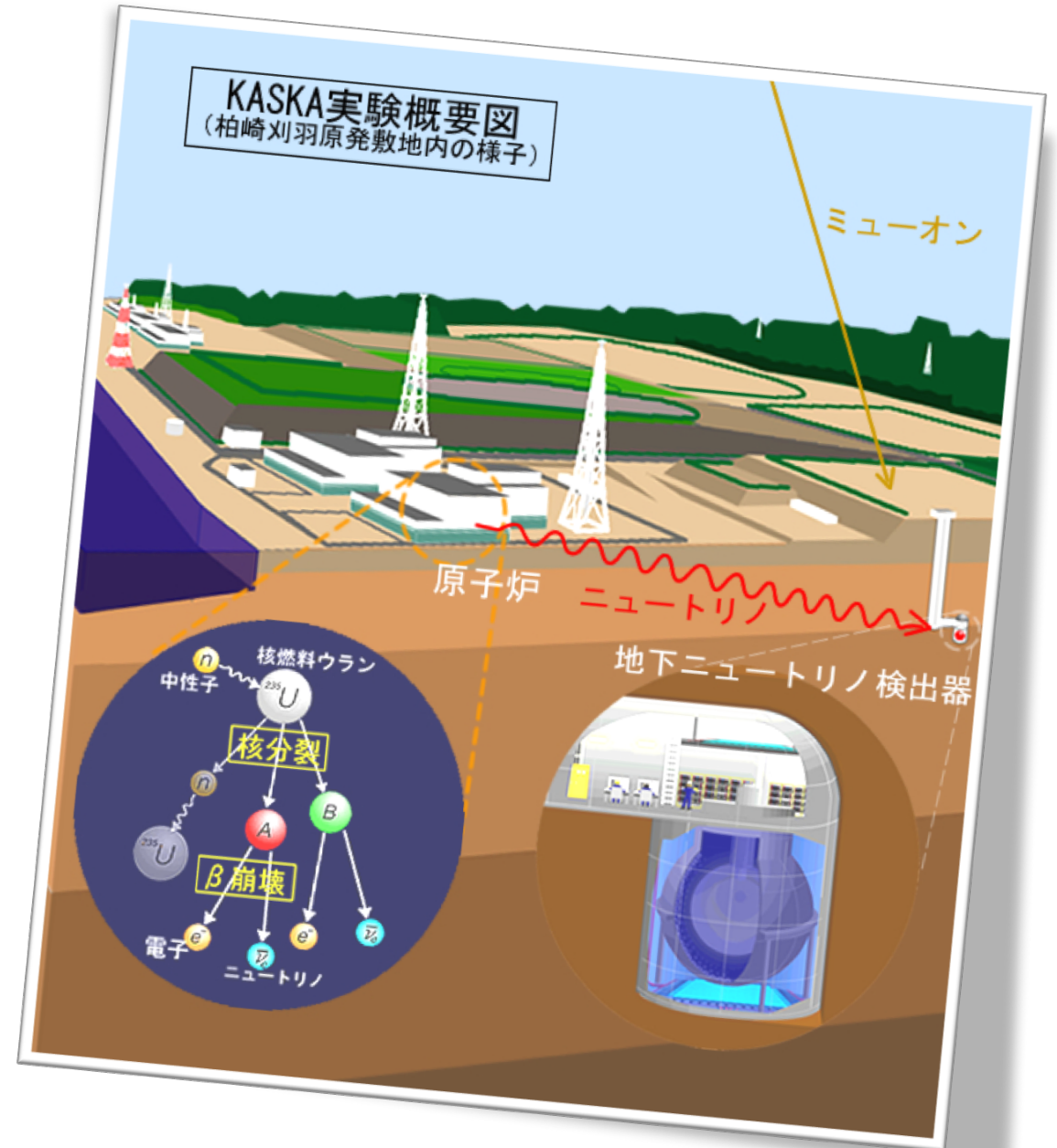
A Precise Measurement of  
 $\sin^2 2\theta_{13}$   
using Japanese Reactors



F.Suekane  
RCNS, Tohoku Univ. Japan

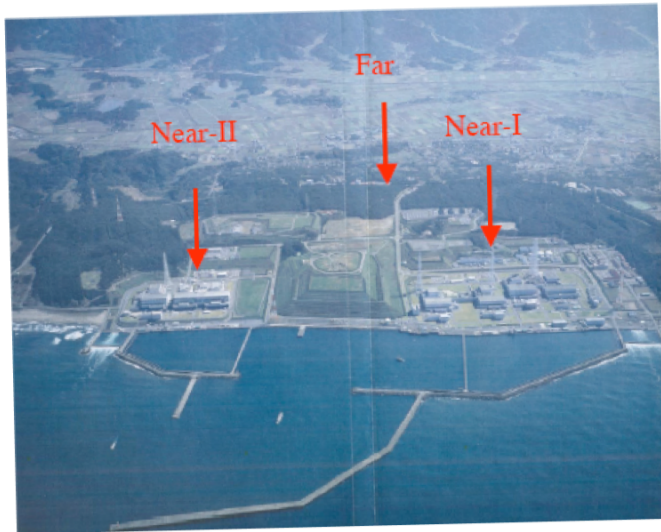
@NOON03  
Kanazawa, Japan

2002.2.10

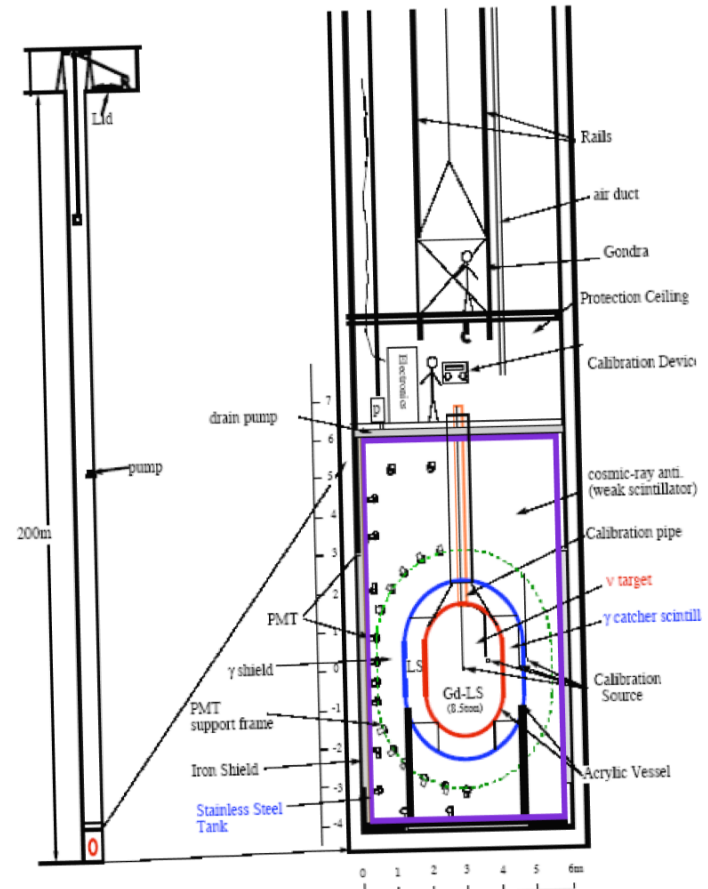


# KASA (Kashiwasaki, Japan)

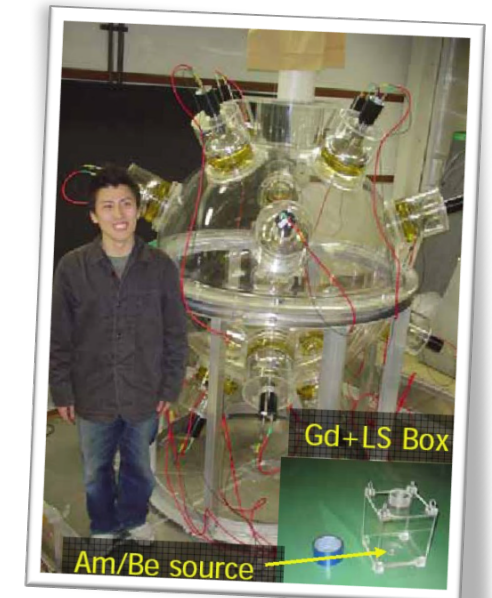
Kashiwazaki-Kariwa  
Nuclear Power Station  
 $P=24.3\text{GW}_{\text{th}}$



3 detectors



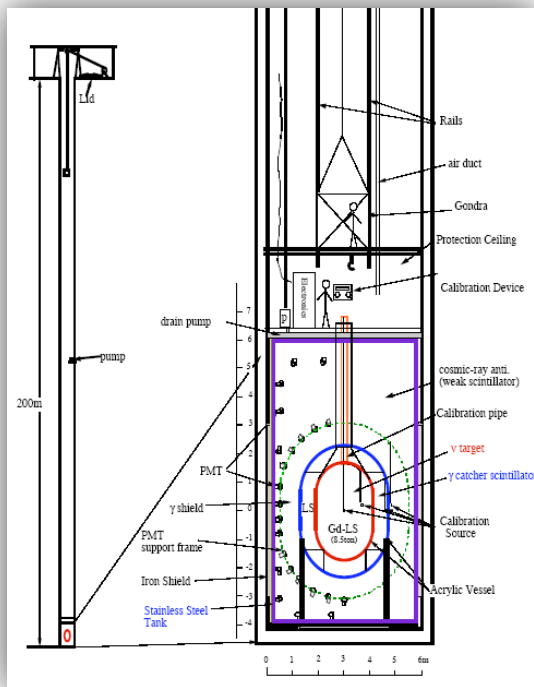
- 2002: KASKA Group was formed
- 2003: Conceptual Design  
Solving Site Availability Issues
- 2004: Detector R&D,
- 2005: Detector R&D
- 2006: Construction
- 2007: Construction
- 2008: Start Data Taking
- 2009: Initial results
- 2010:
- 2011: Final results



Project abandoned (2009) and merged with Double Chooz

# 2011 Earthquake

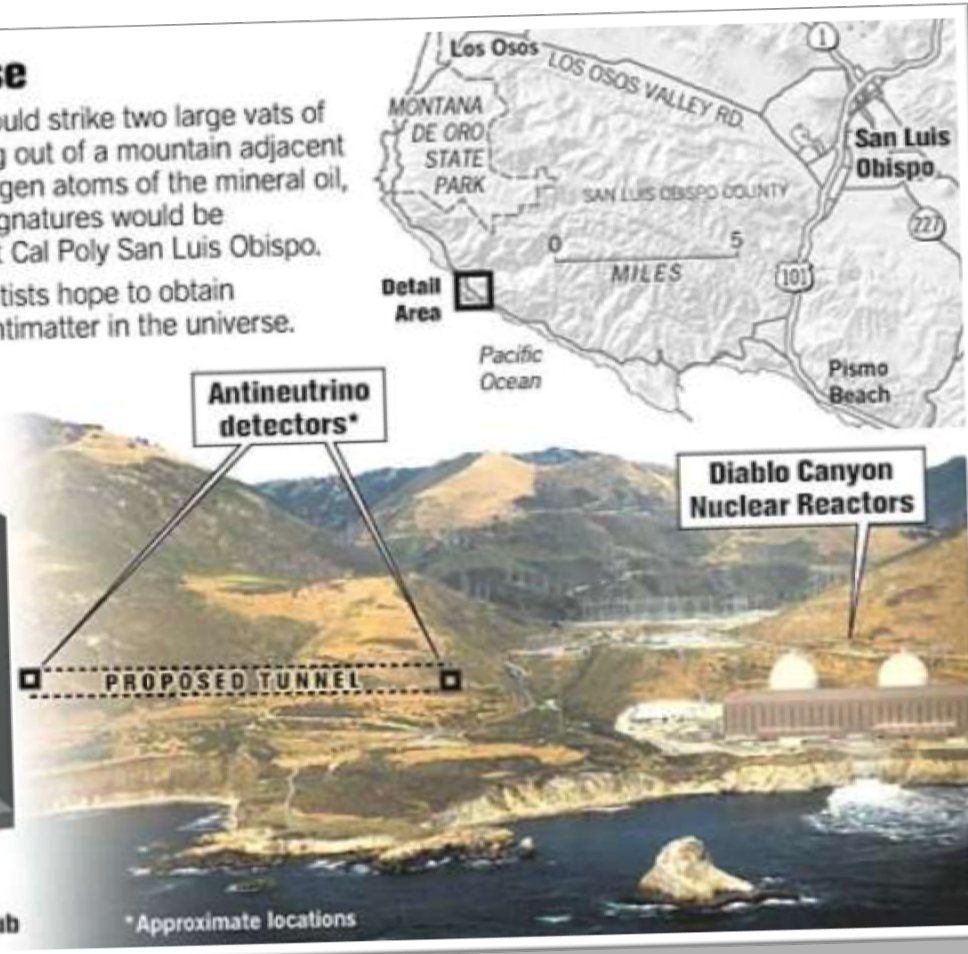
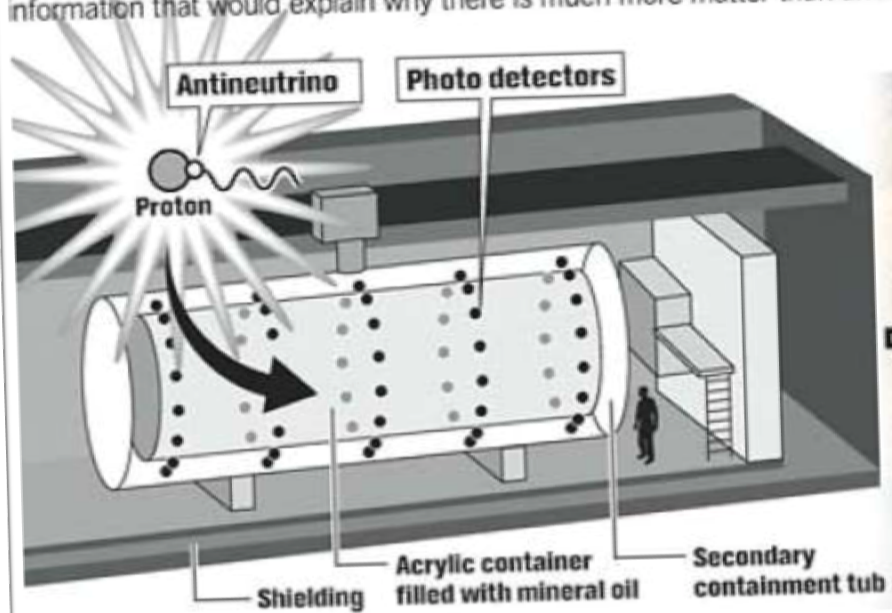
## Major shutdown of Japanese reactors



# Diablo Canyon (California, USA)

## Looking for the building blocks of the universe

In the proposed project, antineutrinos emitted from Diablo's reactor would strike two large vats of mineral oil located at opposite ends of a 1.5- to 2-mile-long tunnel dug out of a mountain adjacent to the power plant. When the antineutrinos strike protons in the hydrogen atoms of the mineral oil, they would make characteristic "signatures," indicating mass. These signatures would be registered by photo detectors and transmitted automatically to a lab at Cal Poly San Luis Obispo. By noting the way the neutrinos "oscillate" between mass states, scientists hope to obtain information that would explain why there is much more matter than antimatter in the universe.



Canyon located at 1.8 km from the reactor cores

Detector hall would have been accessible with modest civil engineering...

Nuclear power plant company under pressure by environmentalists

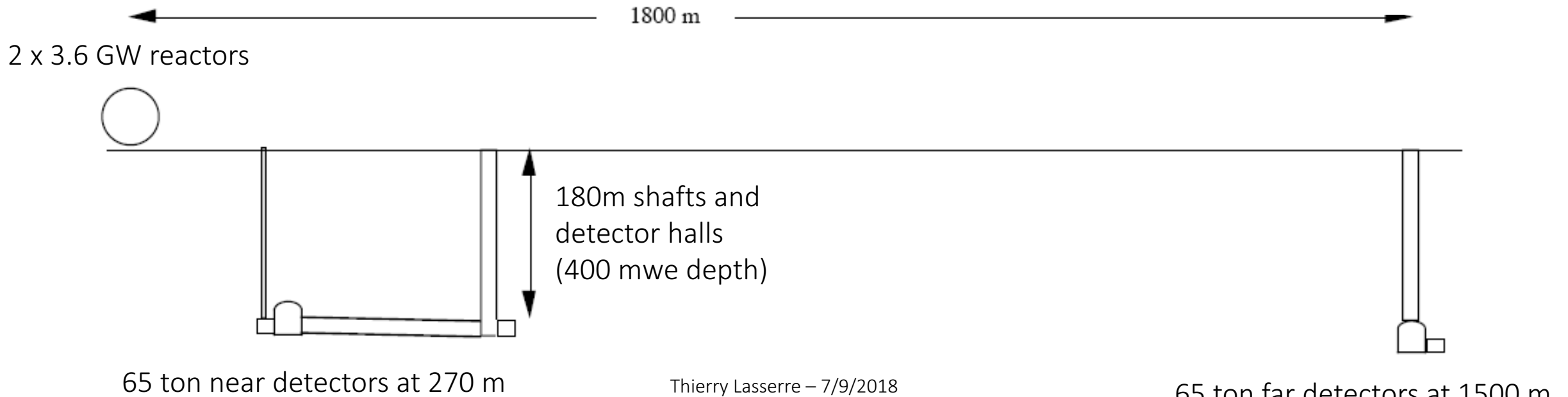
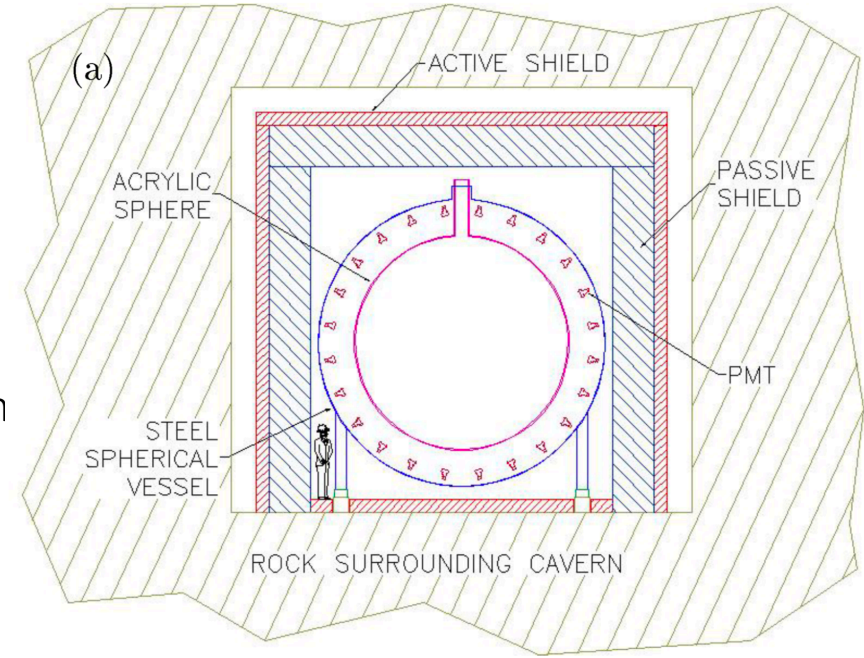
Project abandoned (2004) and somehow realized in China



# Braidwood (Illinois, USA)



- 2 x 3.6 GW reactors
- 180m shafts and halls at 450 mwe depth
- Detector:
  - 2 x 65 ton near detectors at 270 / 1510 n
  - **Spherical shape**
  - **No-gadolinium**



# Angra (Brazil)

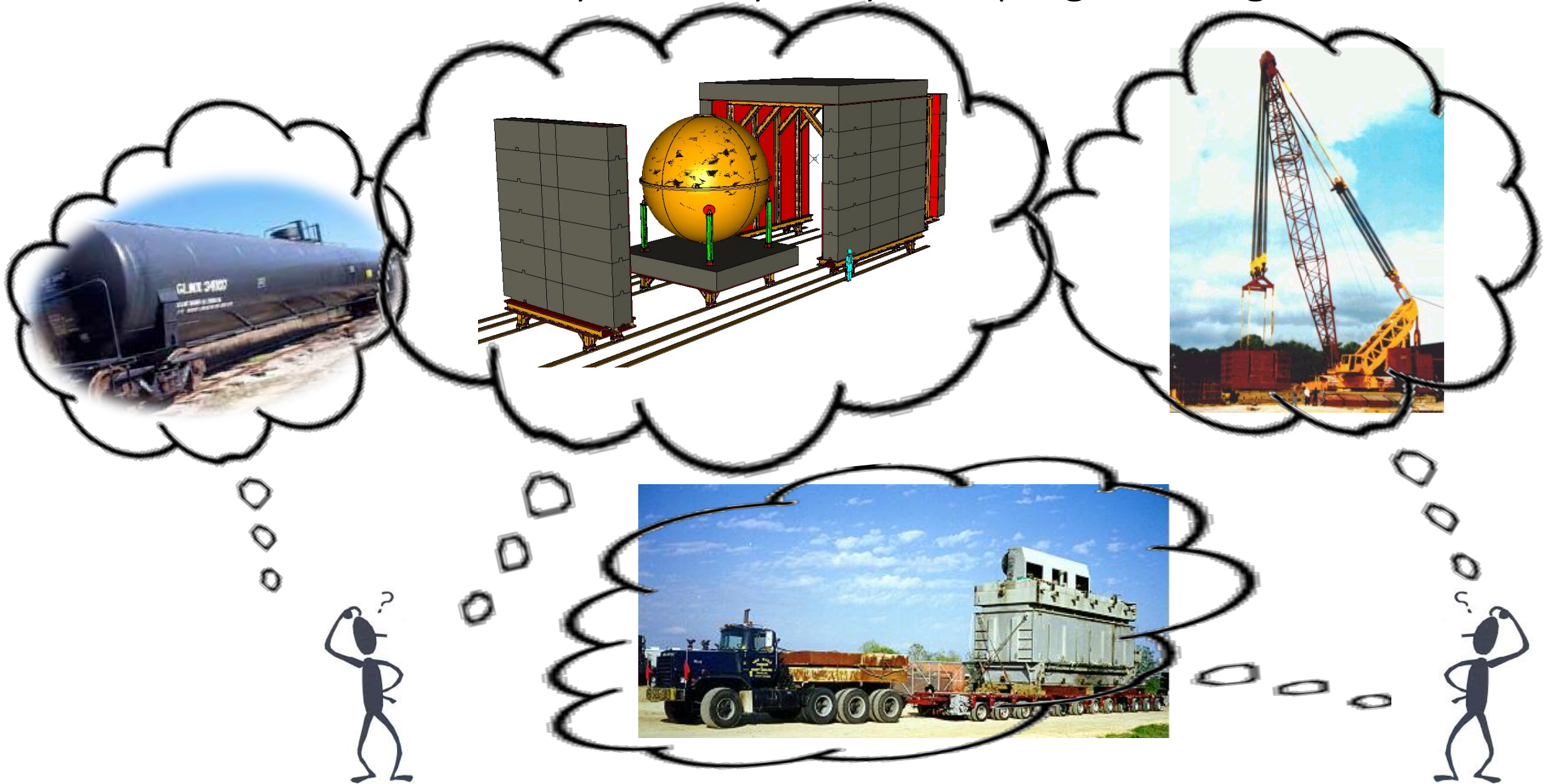


$\theta_{13}$  project too big to be supported by Brazil

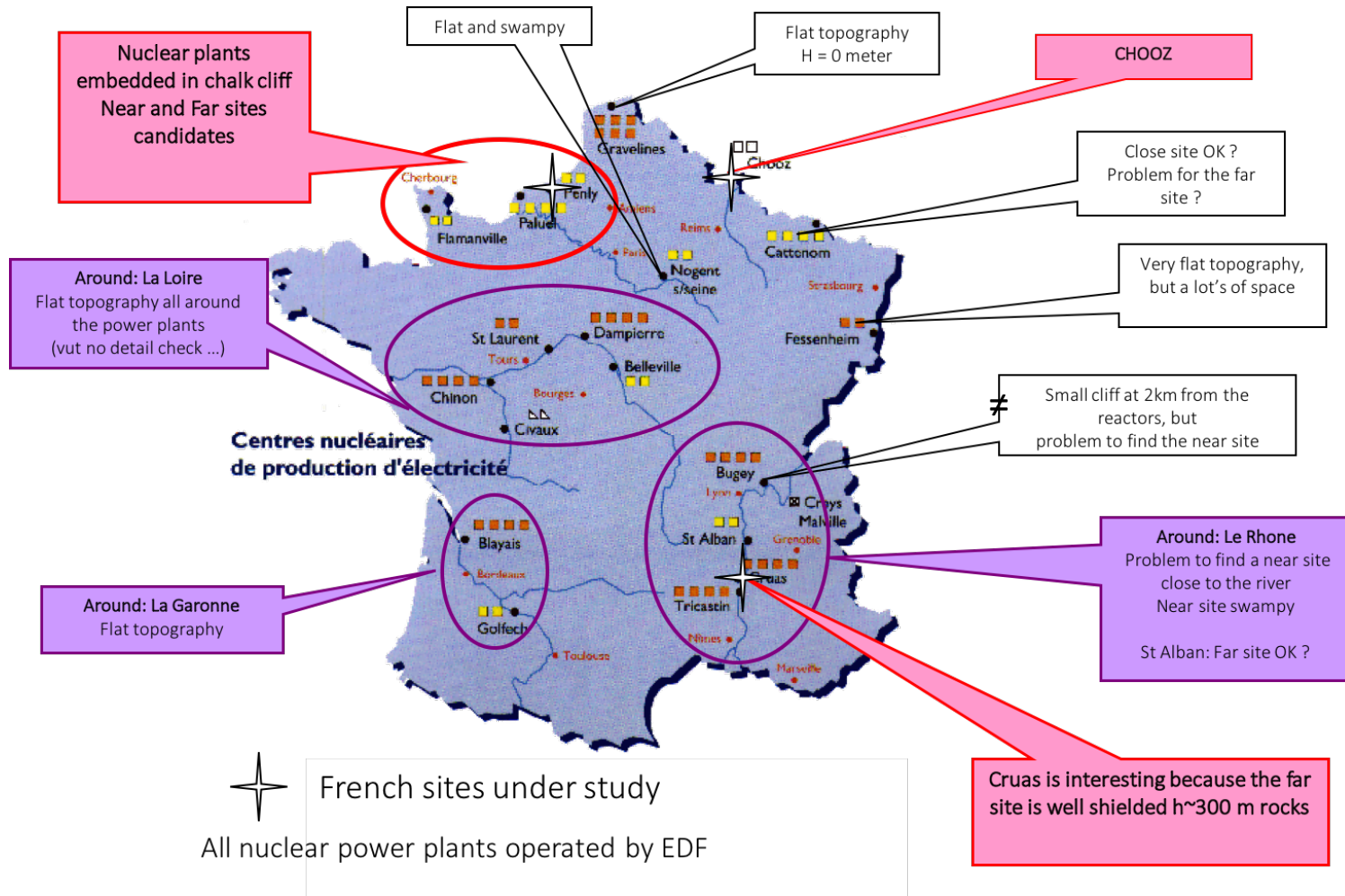


Led to a new reactor neutrino facility (non-proliferation, Connie, ...)

# Braidwood / Diablo Canyon / Daya Bay: Swaping/Moving detectors...

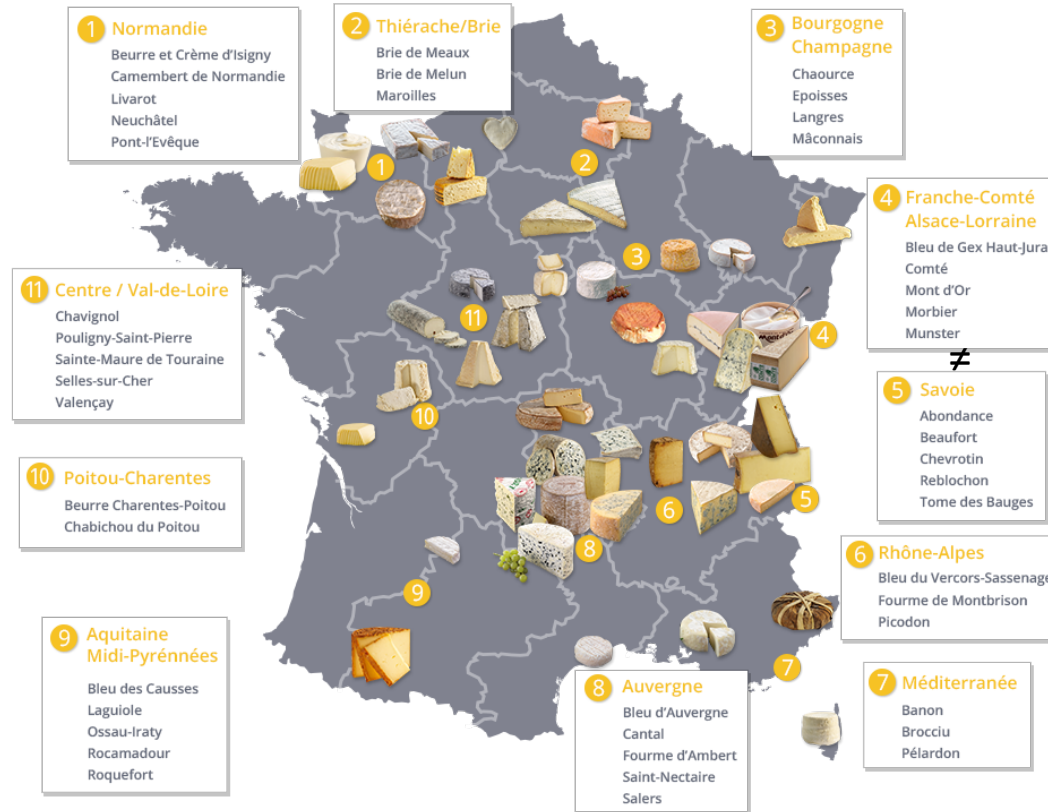


# Site Investigations in France





# Site Investigations in France



≠



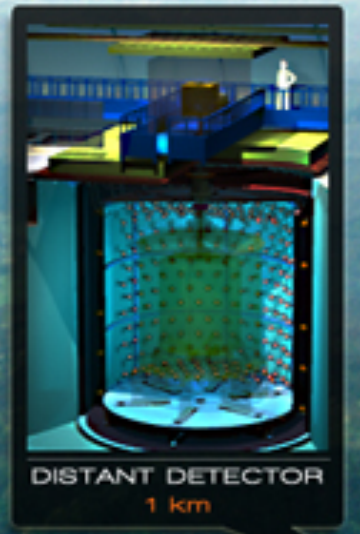
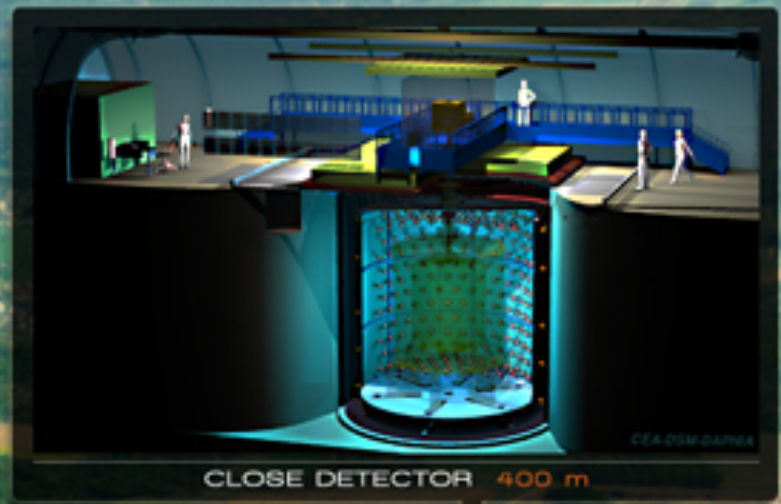
1 type of nuclear reactor  
100 type of cheese

1 type of cheese  
100 type of nuclear reactors

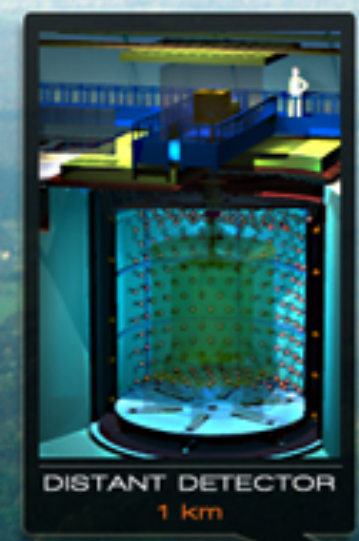
Les gendarmes de Tricastin...



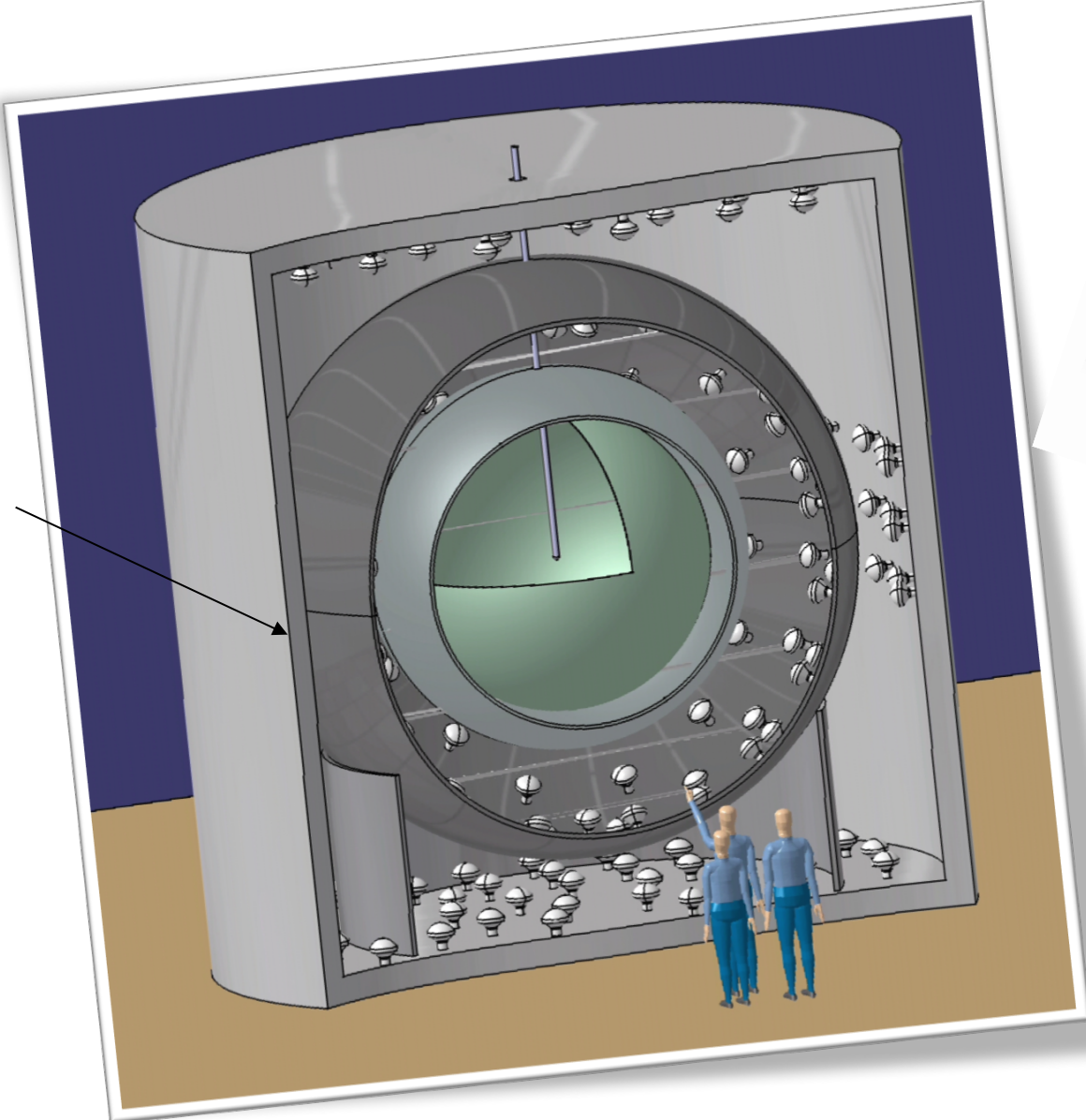
# First Come back to Chooz (2003)



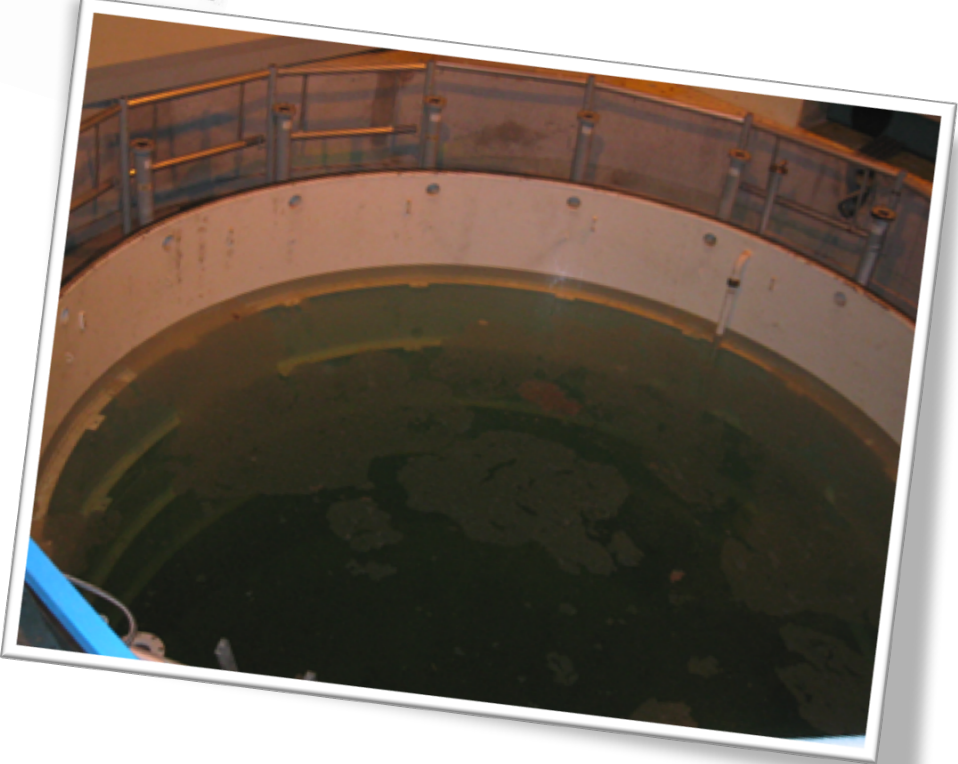
# First Come back to Chooz (2003)



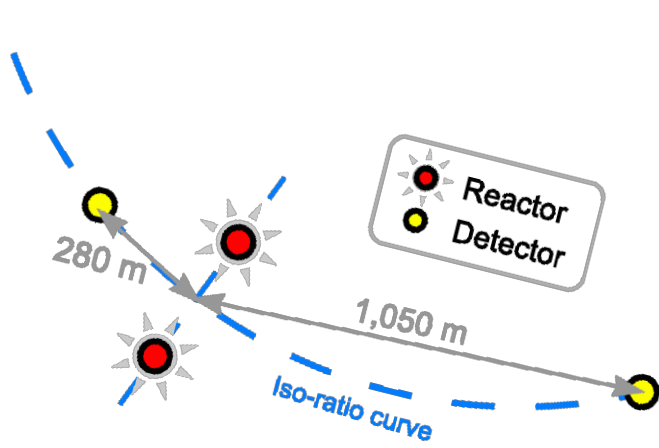
# First Double Chooz – spherical- Design - 2003



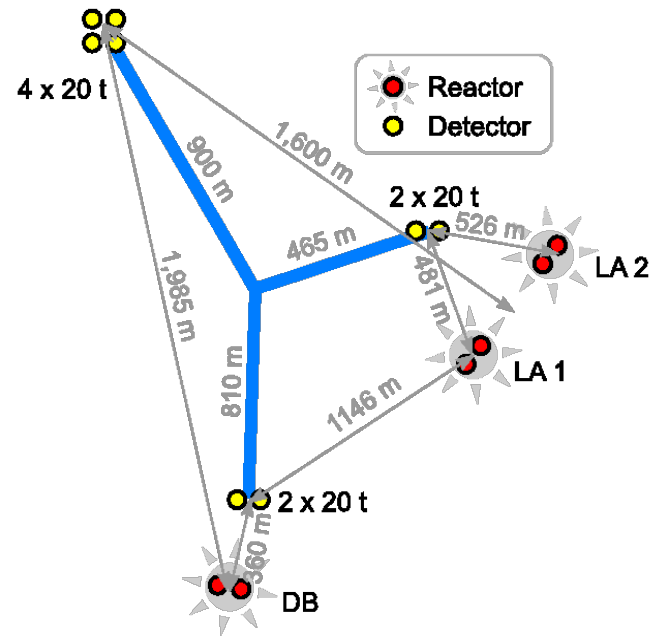
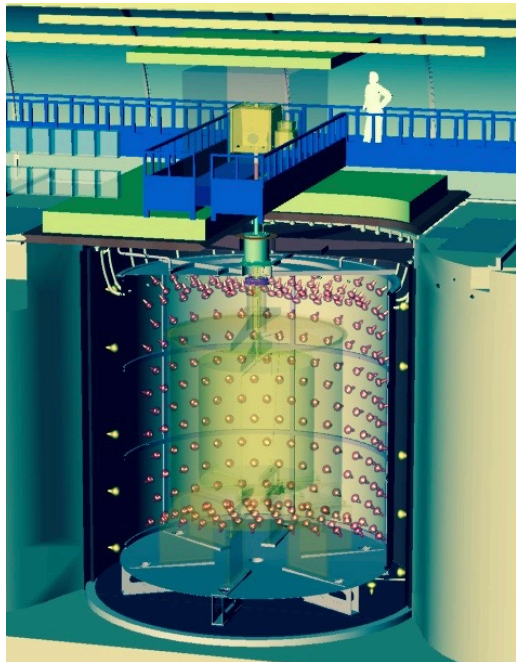
Existing CHOOZ tank



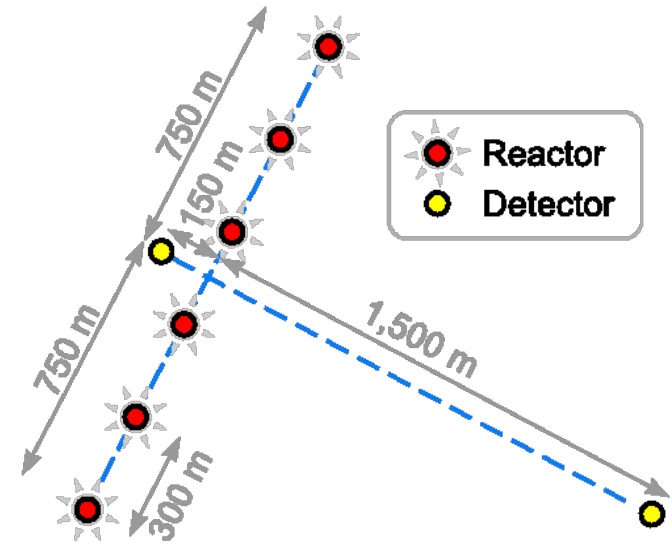
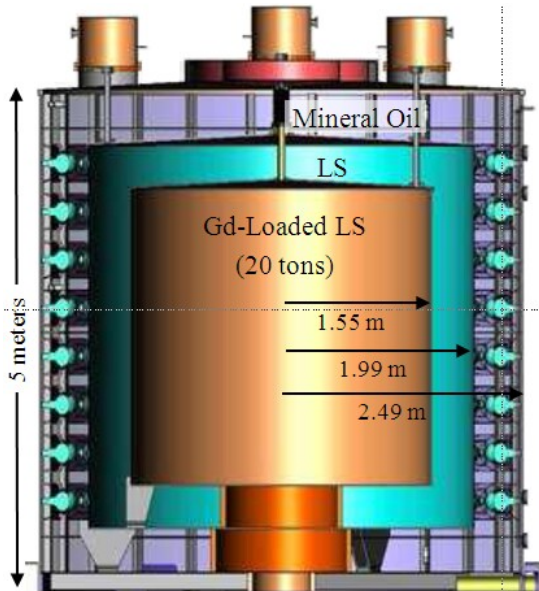
# 12 identical – cylindrical – detectors...



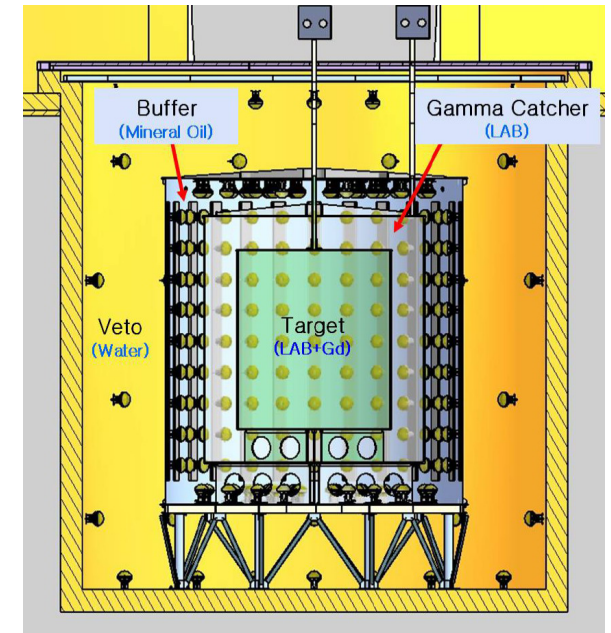
2 detectors Gd-volume: 20 m<sup>3</sup>



8 detectors Gd-volume: 200 m<sup>3</sup>



2 detectors, Gd-volume: 40 m<sup>3</sup>



# Double Chooz

First Double Chooz result in 2011, illustrated in The Big-Bang Theory (CBS)



Started in 2002

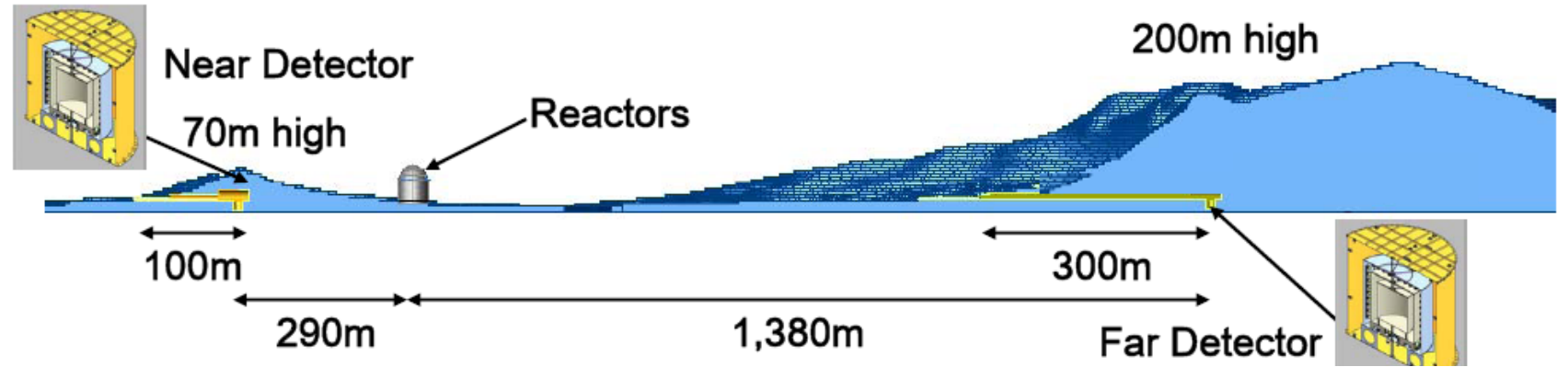
Approved in 2004

Far detector data taking in 2010

(Late) Near laboratory excavation in 2010-12...

Near detector data taking in 2014

# RENO



A Korean-only effort

Starting with (very) little expertise in 2004

About 30 people!

Some cut&paste...

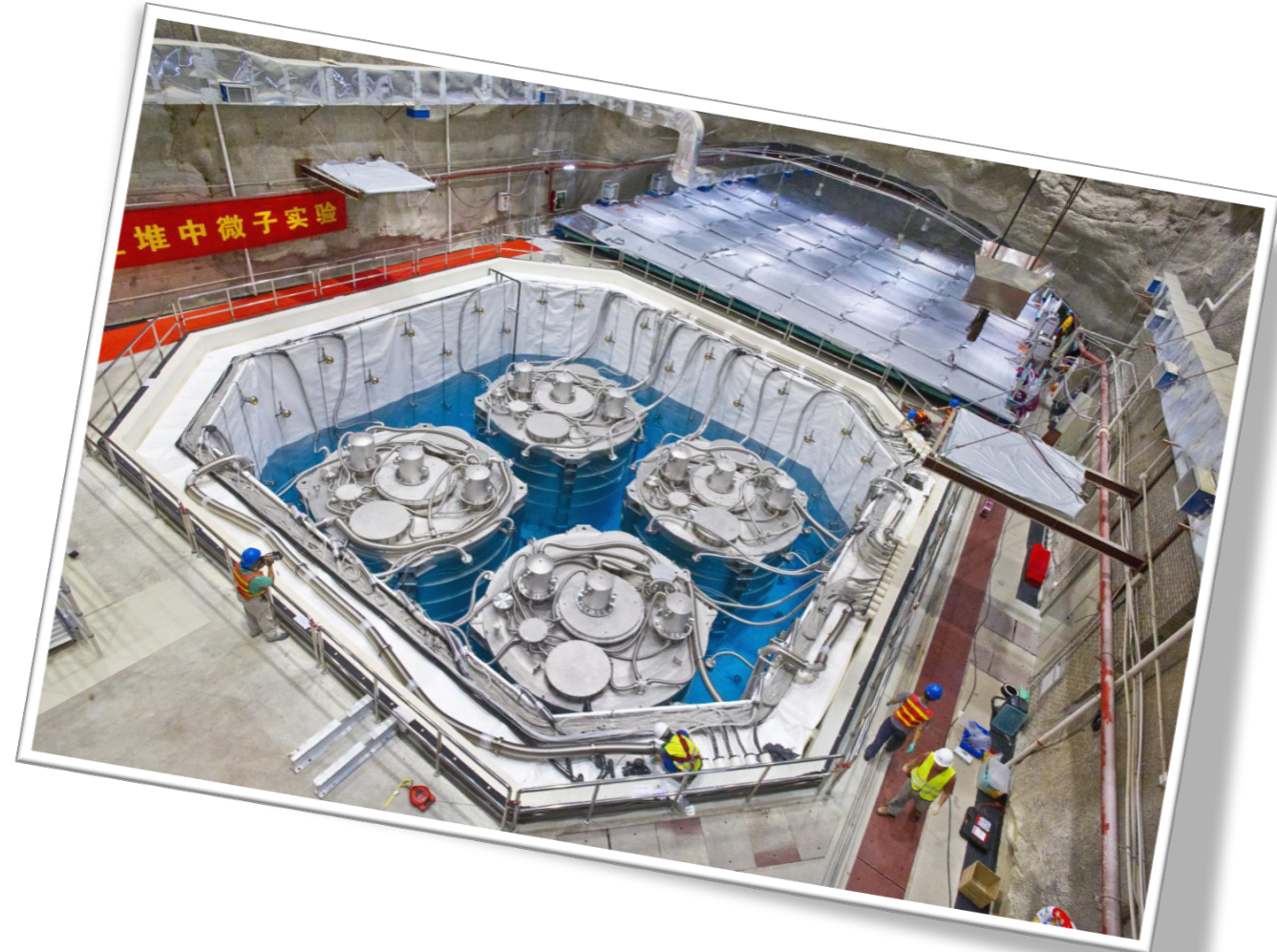
And a great achievement





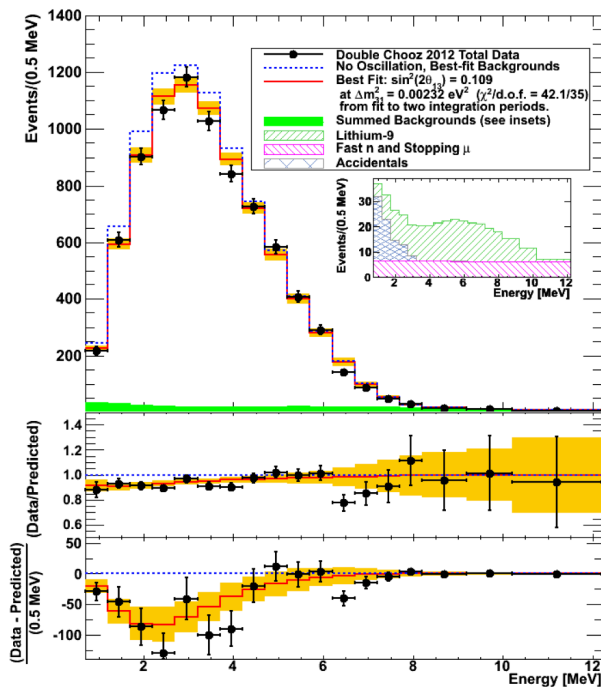
# Daya Bay: A colossal experiment

2006, a high-level political decision leading to the largest international collaboration in basic science in China and the biggest US-China collaboration.



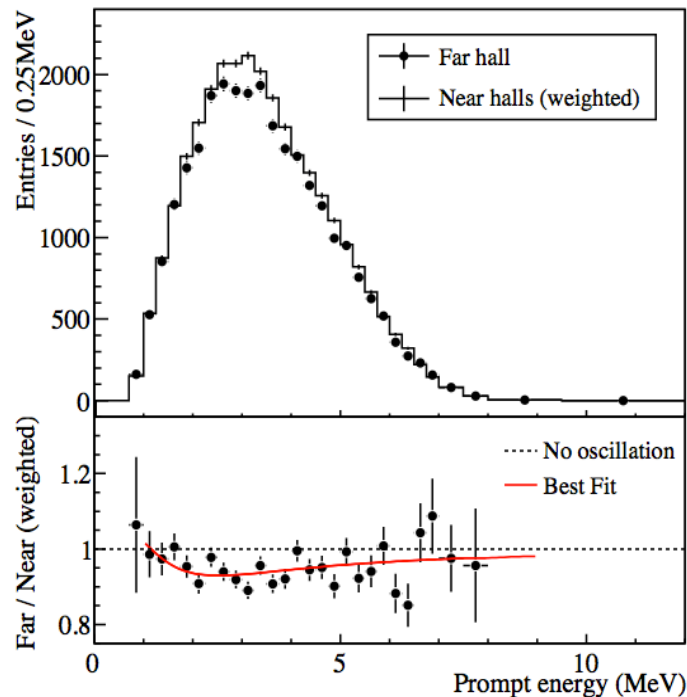
# First Reactor Results (2011-2012)

Double Chooz (Oct. 2011)



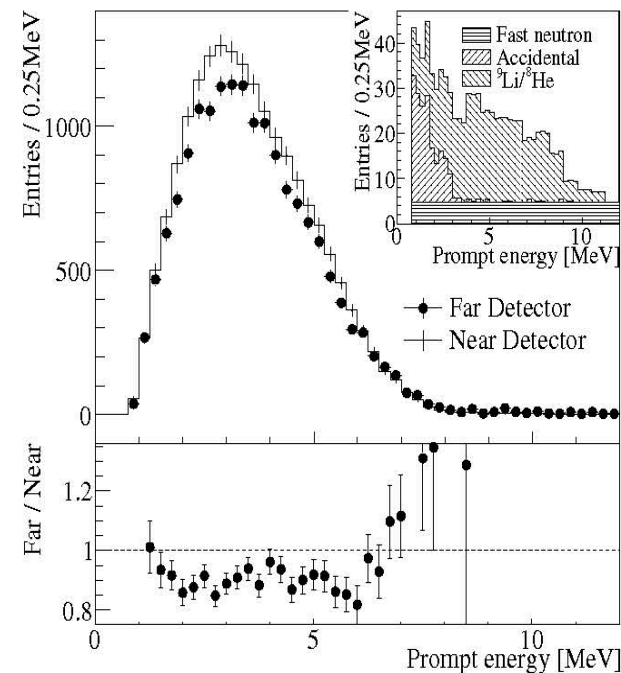
Rate+shape  
98%

Daya Bay (March 2012)



rate only  
 $5 \sigma$   
Leading precision - 15% error

RENO (April 2012)

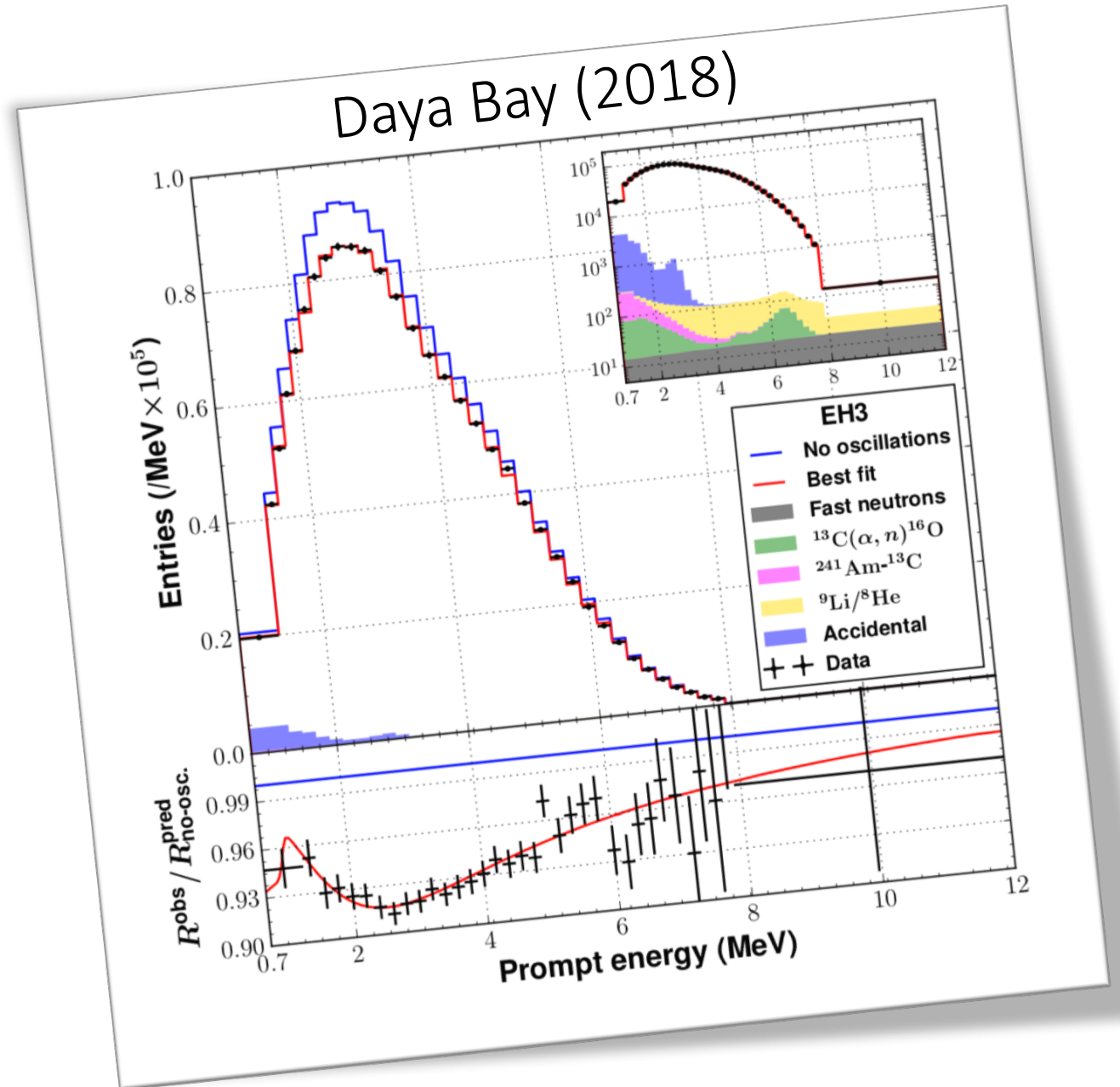


Rate only  
Almost  $5 \sigma$

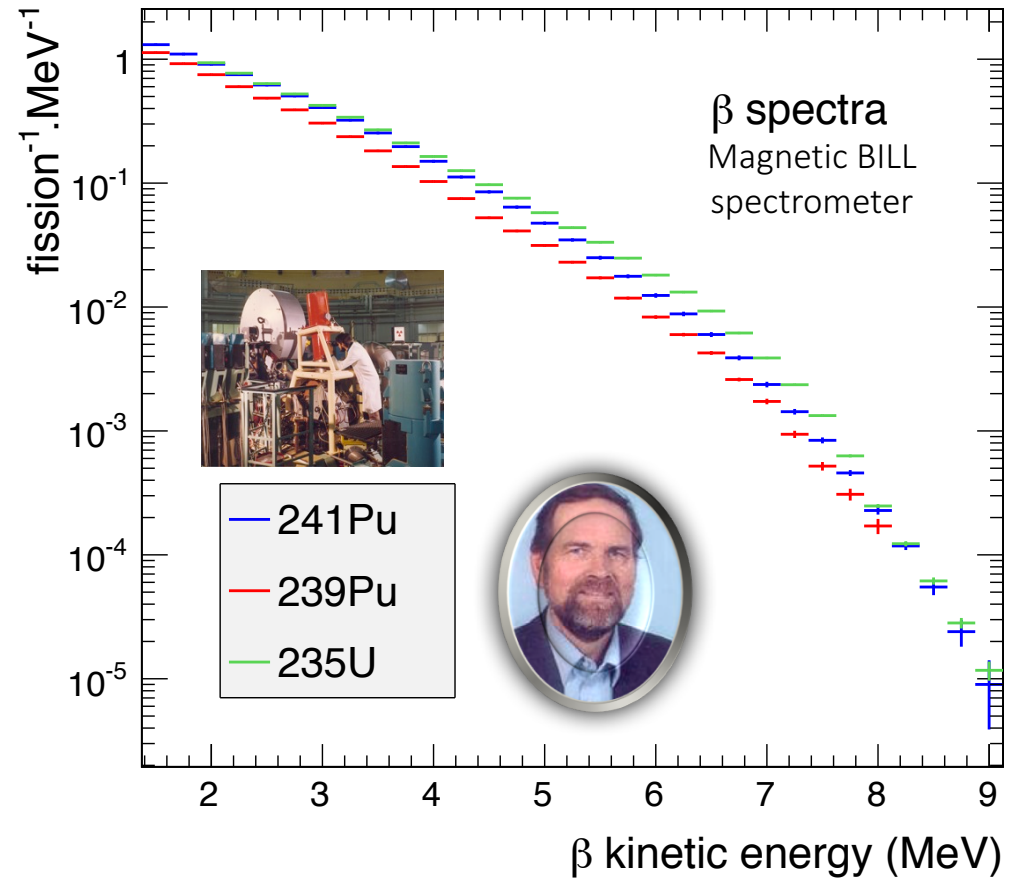
Reactor experiments  
provide the most  
precise value of  $\theta_{13}$

$\Delta m^2$       1.4 %  
 $\delta m^2$       2.2 %  
 $\sin^2\theta_{13}$     3.8 %  
 $\sin^2\theta_{12}$     4.4 %  
 $\sin^2\theta_{23}$     ~ 5 %

(1  $\sigma$  uncertainty)

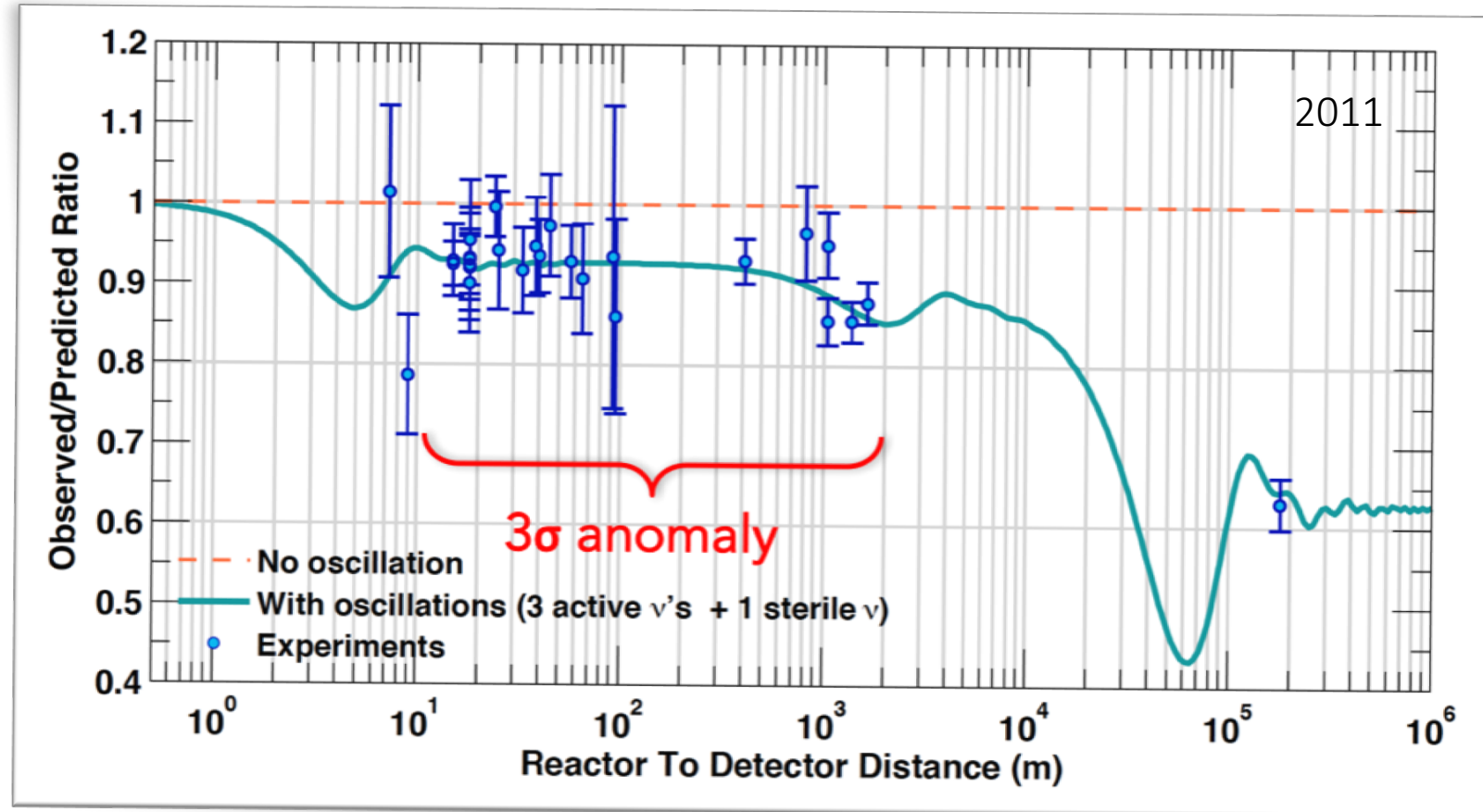
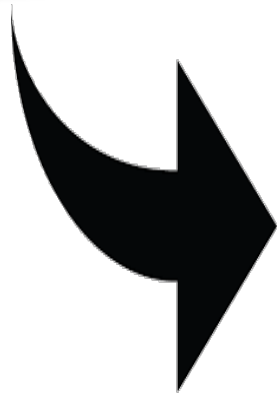
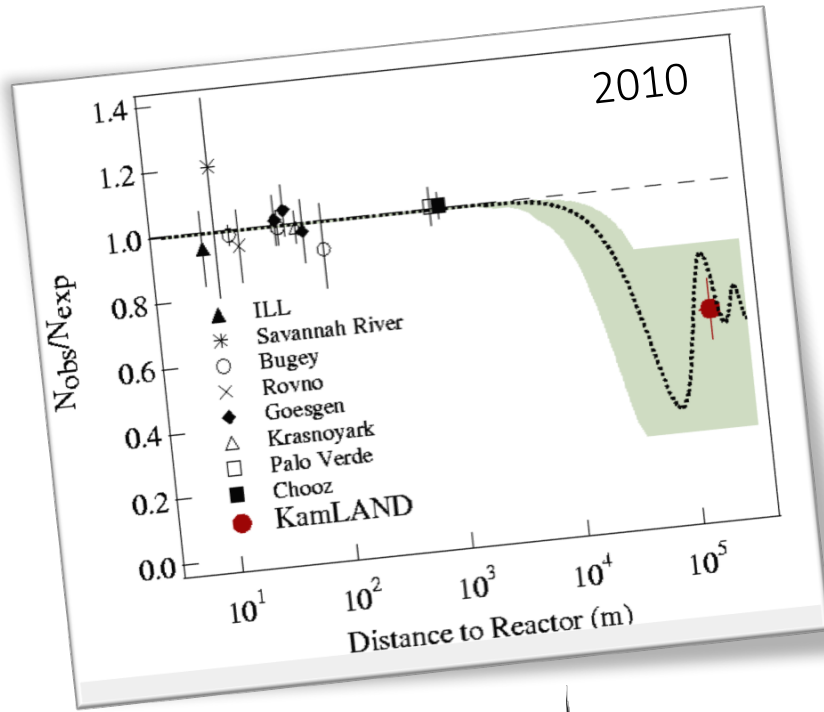


# Reactor Neutrino Flux and Spectra

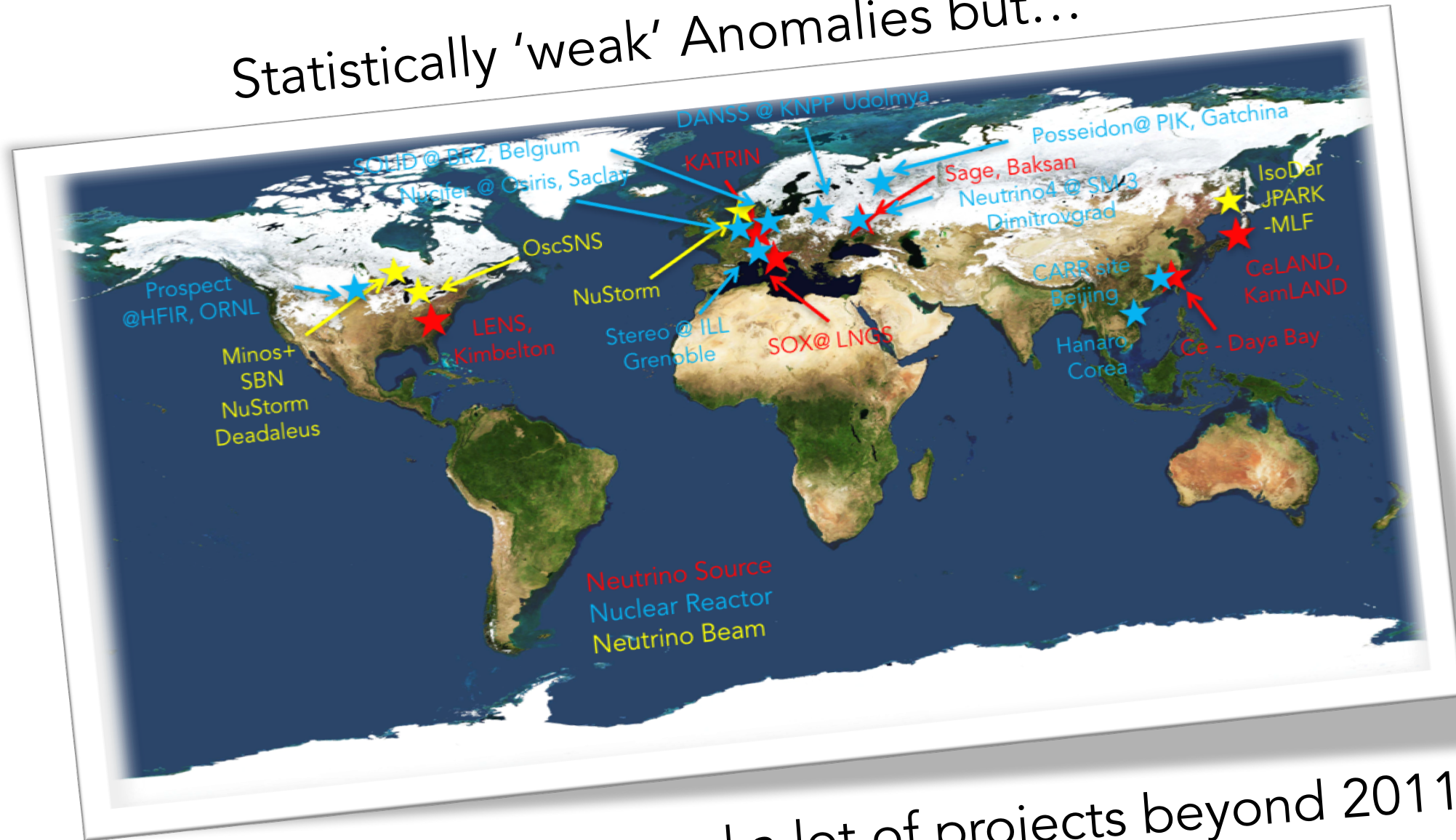


2011: Reevaluation of the  $e - \nu$  conversion procedure – Flux reevaluated at + 3.5%!  
(motivation: interpret Double Chooz results, without near detector)

# Reactor Antineutrino Anomaly - 2011



# Statistically 'weak' Anomalies but...



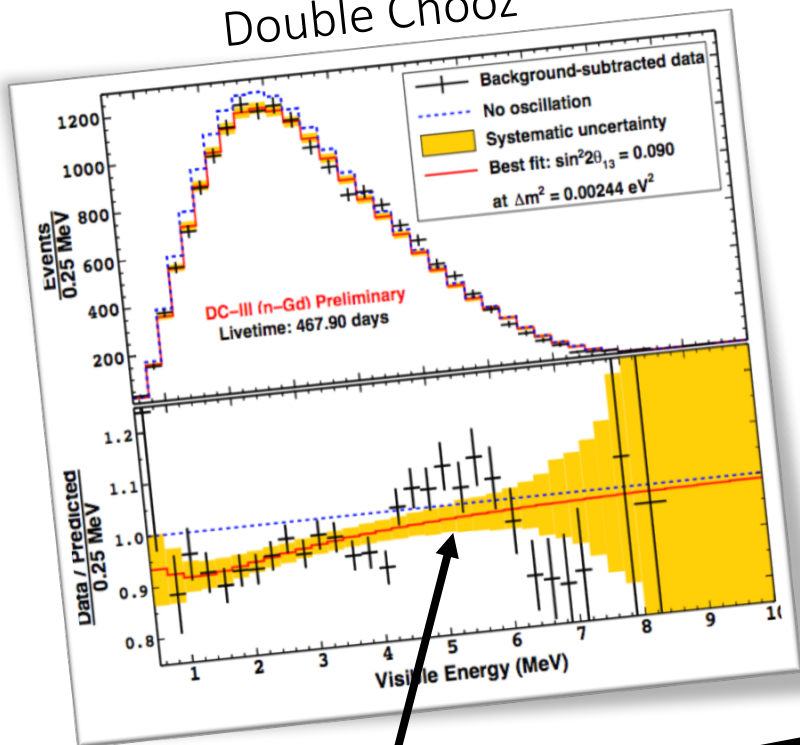
Triggered a lot of projects beyond 2011

A new anomaly (almost) revealed

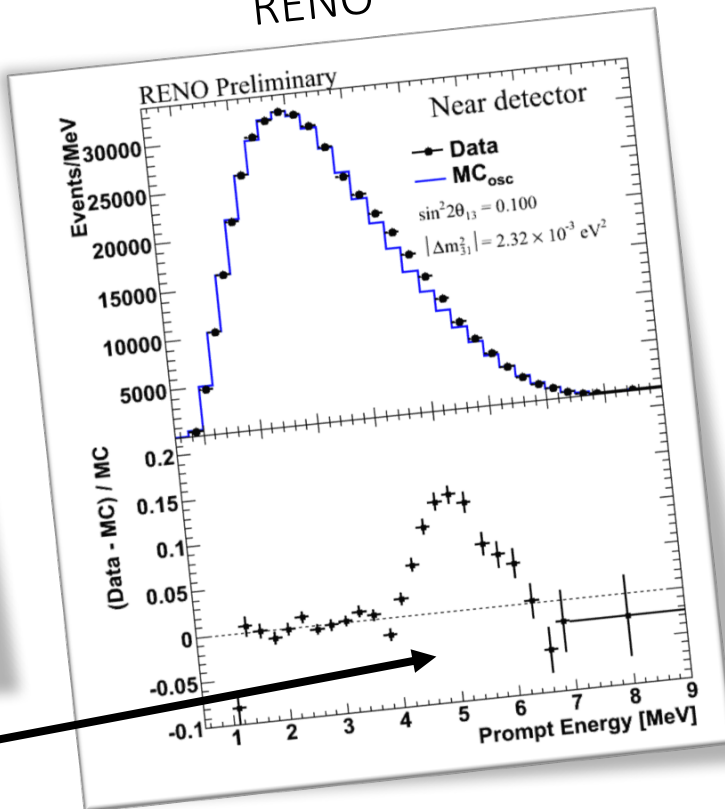
# NEUTRINO 2014

XXVI International Conference on Neutrino Physics and Astrophysics

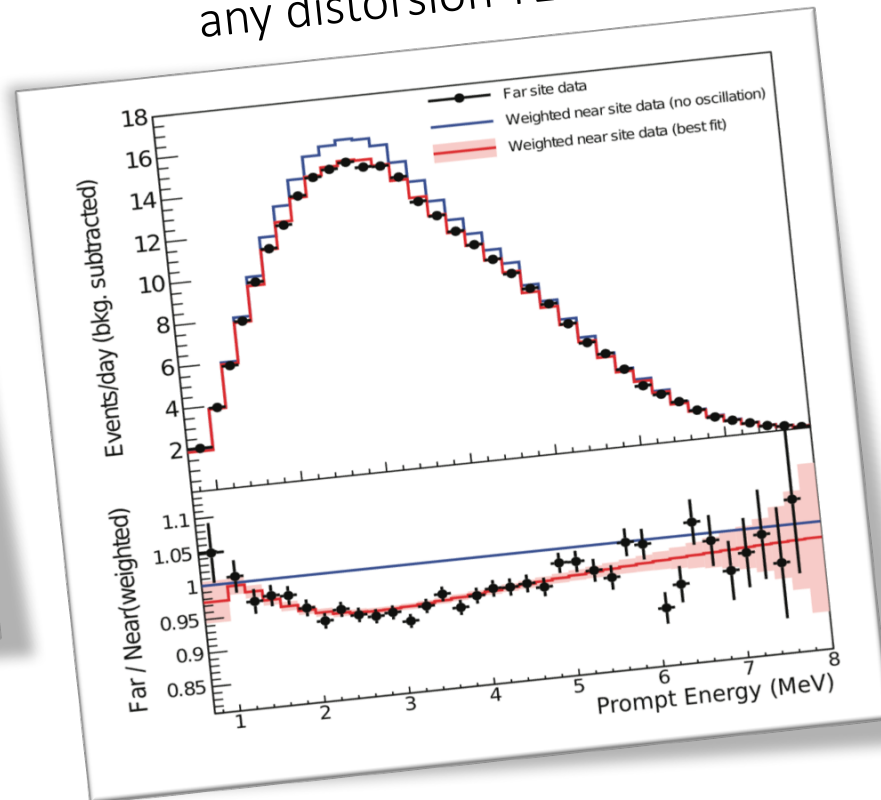
Double Chooz



RENO

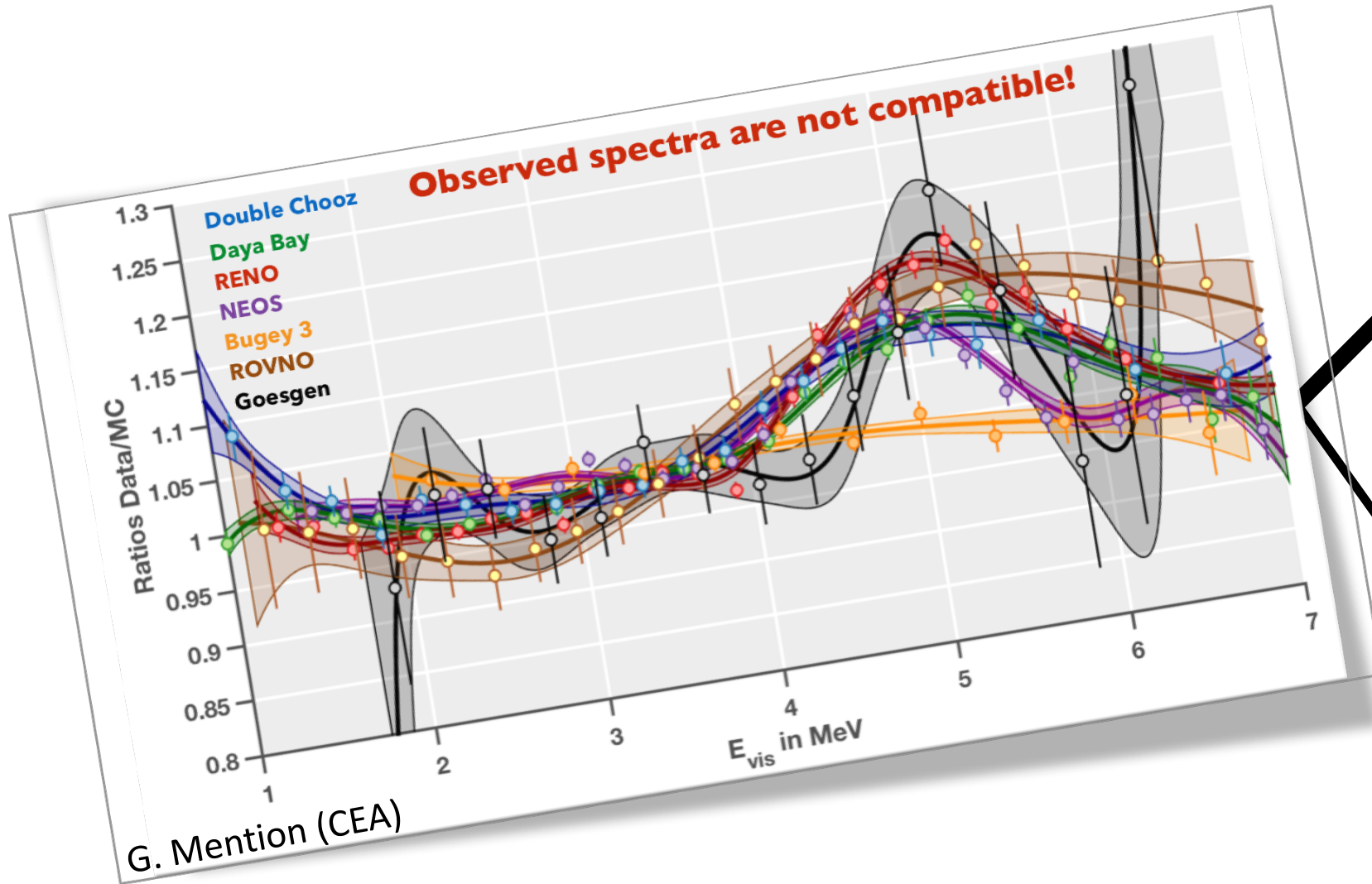


BUT Daya Bay did not report any distortion YET...



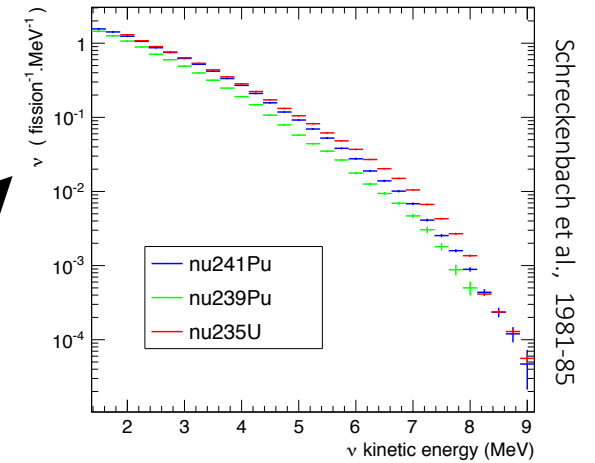
So-called 5 MeV bump

# 4-6 MeV spectral distortion still unexplained...



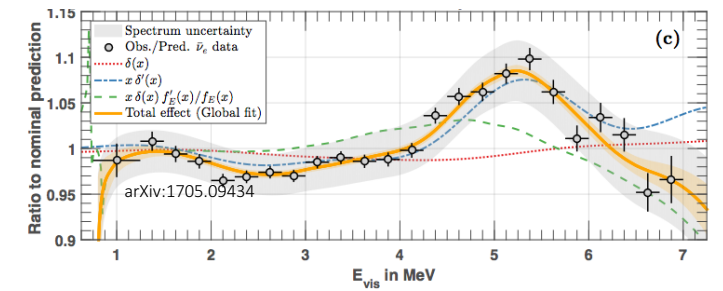
## Reactor $\nu$ -spectra?

Bias? Underestimated systematics?



## Detector response?

1% E-scale non-linearity?





# To Hervé de Kerret (1947-2017)

Double-CH $\theta_{13}\theta_{13}Z$  ?

$\theta_{\text{CHOOZ}}$  ?

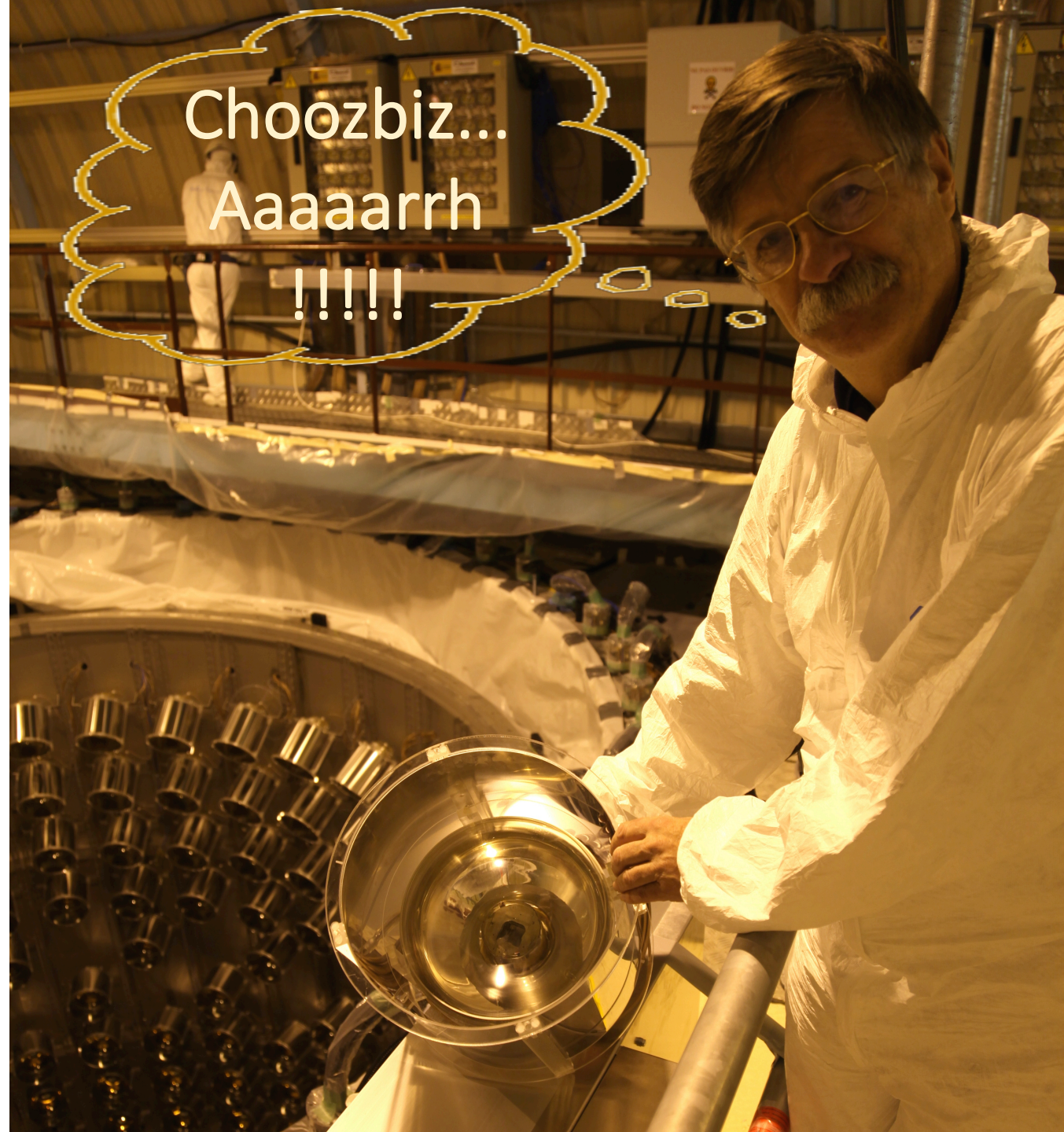
CH $\theta_{13}\theta_{13}Z$ -2D ?

CH $\theta_{13}\theta_{13}Z$  ?

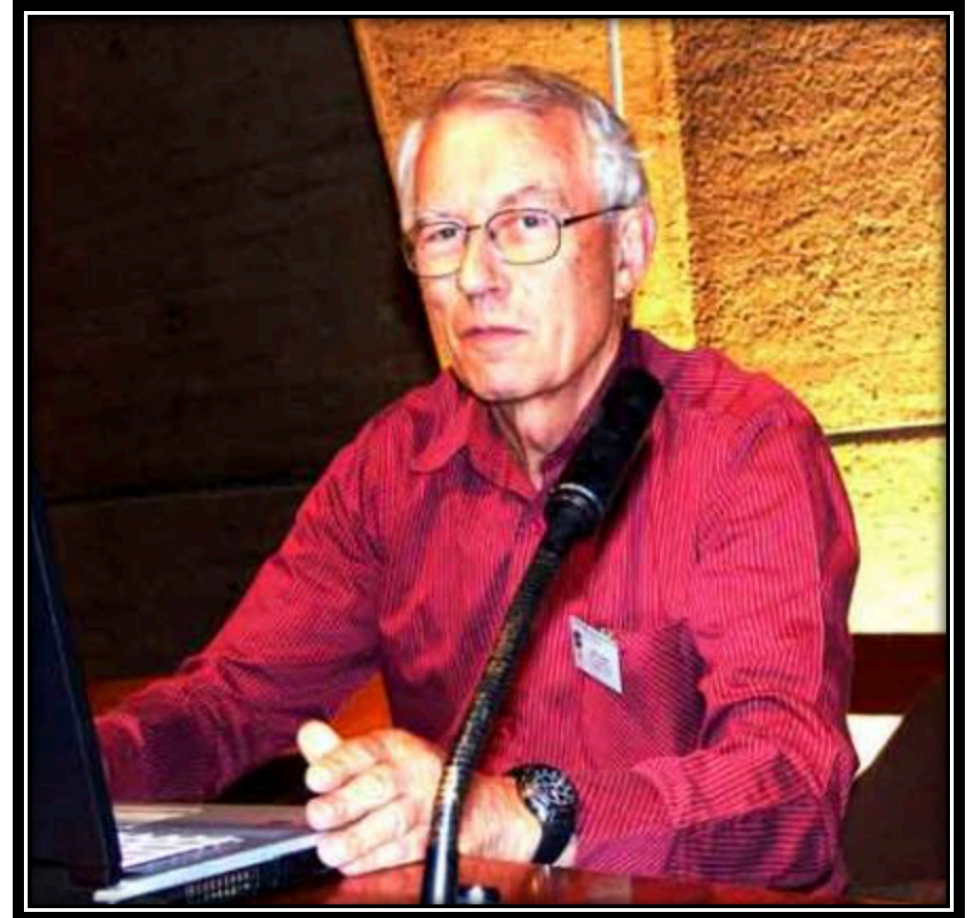
Super-CH $\theta_{13}\theta_{13}Z$  ?

CH $\theta_{13}\theta_{13}Z$ -II

....



# Thank you!



Objet: The Reactor Neutrino Anomaly...

Date: 19 décembre 2010 à 17:27:14 UTC+1

I synthesized all the known measurements at short distances (<Chooz)... You will see that the result of this synthesis are really very surprising....

ALL measurements (except 1) are more than 1 sigma from the LM prediction. The average of the effect is 0.943 and the sigma is 0.03 ...

The options available to us:

- Anomaly in the calculation of the new reactor neutrino spectra

OR

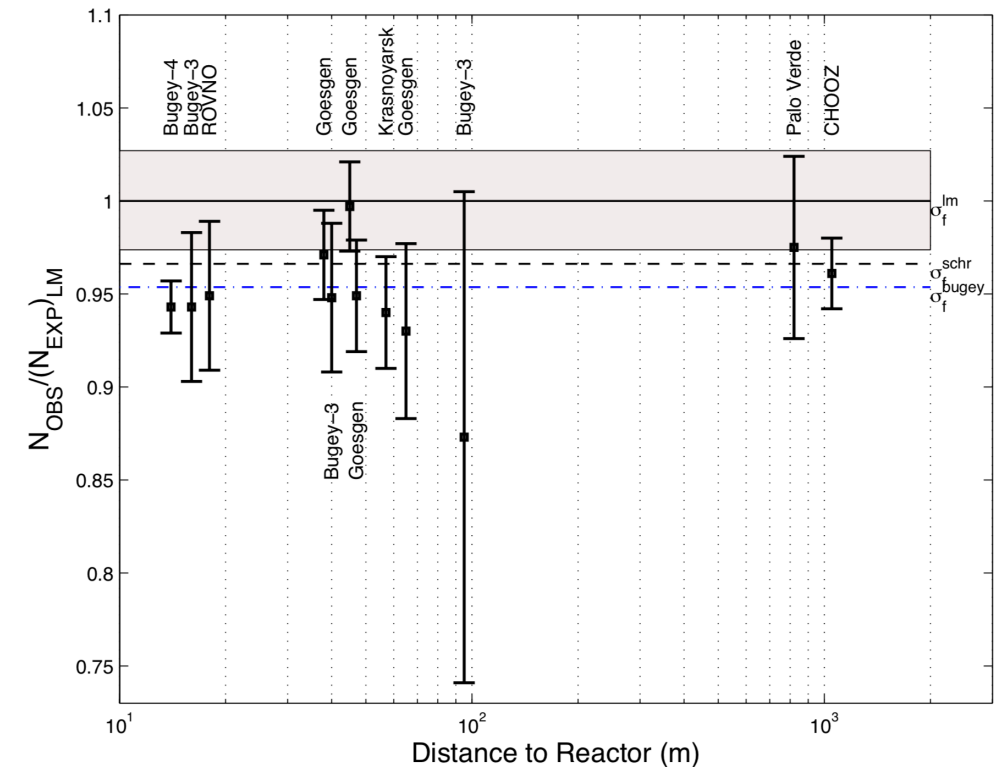
- Anomaly in the measurement of neutrinos

- Case 1: All the experiments underestimate their errors and/or do a bad count, but strange that it is always in the same direction?

- Case 2: Neutrino oscillation but without spectrum distortion to match Bugey3 & Goesgen spectrum ratios, so only possible with massive neutrinos,  $\gg 1$  eV (keV?). This solution is ULTRA-hypothetical but I prefer to list it as a possibility for completeness. An effect of  $1 - 0.943 = 0.057$  would result, at high  $Dm^2_{41}$ , in  $\sin^2(2\theta_{14}) = 0.114$  ... which is so large that it could be incompatible with other neutrino oscillation results (solar, atm, beams, etc.), wouldn't it? Thank you for clarifying this point.

- The neutrinos are still surprising us....

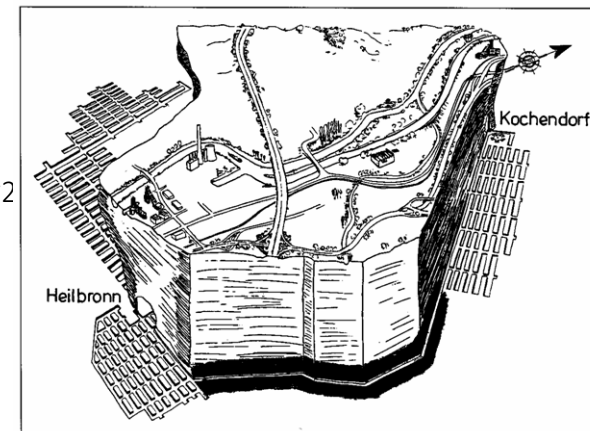
# Reactor Antineutrino Anomaly – 12/2010



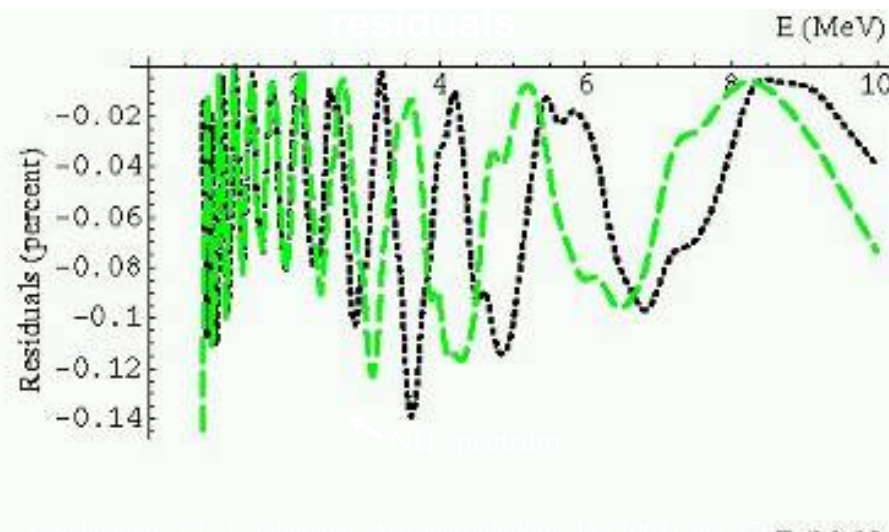
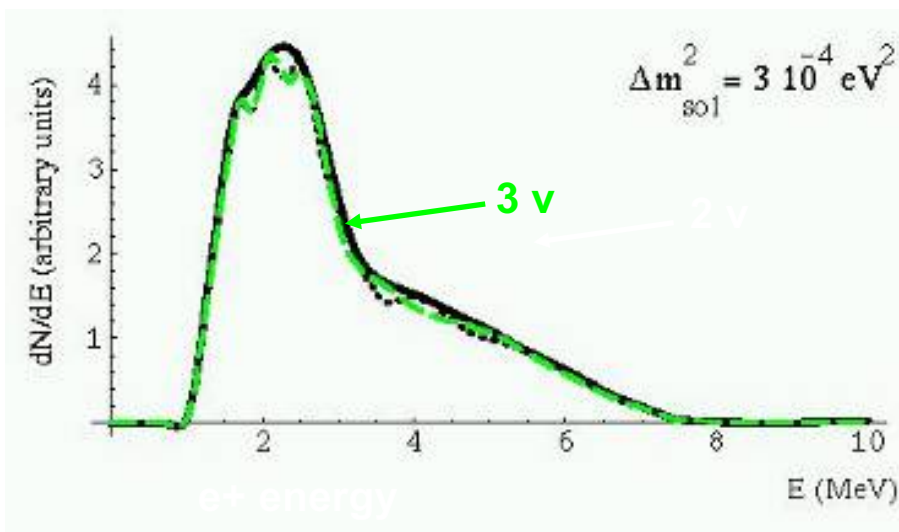


# HLMA: a failed attempt

The Heilbronn salt mine



- Context <2003: KamLAND could not have measured  $\Delta m_{\text{sol}}^2$  if greater than  $2 \cdot 10^{-4} \text{ eV}^2$
- Study of the HLMA: a new reactor neutrino project – relevant only if  $\Delta m^2 > 2 \cdot 10^{-4} \text{ eV}^2$
- By products with a 1 kiloton detector Heilbronn Mine (20 km baseline)
  - constraint on  $(\sin^2 2\theta_{13}) > 0.05$
  - probe of the  $\nu$  mass hierarchy



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \underbrace{2|U_{e3}|^2(1-|U_{e3}|^2)}_{\text{atm}}(1-\cos\Delta_{31}) - \underbrace{\frac{1}{2}(1-|U_{e3}|^2)^2 \sin^2(2\theta_{12})(1-\cos\Delta_{21})}_{\text{solar}} \\
 + \underbrace{2|U_{e3}|^2(1-|U_{e3}|^2)}_{\text{interference}} \sin/\cos^2(\theta_{12})(\cos(\Delta_{31}-\Delta_{21})-\cos\Delta_{31}) \quad \Delta_{ij} = \frac{\Delta m_{ij}^2 L}{2E}$$